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# OPTIMAL SHAPE PARAMETER FOR MESHLESS SOLUTION OF THE 2D HELMHOLTZ EQUATION



PARÁMETRO ÓPTIMO DE FORMA PARA LA SOLUCIÓN DE LA ECUACIÓN DE HELMHOLTZ 2D EN MODELO SIN MALLA

Mauricio-A, Londoñoª\*; Hebert, Montegranarioª.

#### ABSTRACT

The solution of the Helmholtz equation is a fundamental step in frequency domain seismic imaging. This paper deals with a numerical study of solutions for 2D Helmholtz equation using a Gaussian radial basis function-generated finite difference scheme (RBFFD). We analyze the behavior of the local truncation error in approximating partial derivatives of the 2D Helmholtz equation solutions when the shape parameter of RBF varies. For discretization, we performed, by means of a classical numerical dispersion analysis with plane waves, a minimization of the error function to obtain local and adaptive near optimal shape parameters according to the local wavelength of the required solution. In particular, the method is applied to obtain a simple and accurate solver by using stencils which seven nodes on hexagonal regular grids, wich mitigate pollution-effects. We validated numerically that the stability and isotropy are enhanced with respect to Cartesian grids. Our method is tested with standard case studies and velocity models, showing similar or better accuracy than finite difference and finite element methods. This is an efficient way for interacting with inverse and imaging problems such as Full Wave Inversion

#### RESUMEN

La solución de la ecuación de Helmholtz es una parte fundamental en la modelación sísmica en el dominio de la frecuencia. Este artículo realiza un análisis numérico de las soluciones de la ecuación de Helmholtz 2D utilizando un esquema de diferencias finitas (RBFFD), generado por funciones gaussianas de Base radial. Se analiza el comportamiento del error de truncamiento local al aproximar las derivadas parciales de las soluciones de la ecuación de Helmholtz 2D cuando varía el parámetro de forma de la RBF. Para la discretización, hemos realizado, mediante un análisis de dispersión clásico con ondas planas, una optimización de la función de error para obtener valores locales y adaptativos del parámetro de forma de acuerdo con la longitud de onda local de la solución requerida. En particular, el método se aplica para obtener un programa sencillo y óptimo usando plantillas de siete nodos sobre mallas regulares hexagonales, que mitigan el efecto polución. Se comprueba numéricamente que la estabilidad e isotropía son mejoradas con respecto a las mallas cartesianas. Nuestro método es probado con casos de estudio y modelos de velocidad estándar, mostrando exactitud similar o mejor que los métodos de diferencias o elemento finitos. Esta es una manera eficiente de interactuar con problemas inverso y de imagen tales como la inversión de onda completa

#### **KEYWORDS / PALABRAS CLAVE**

**RBF-FD | Helmholtz equation | Shape parameter | pollution effect.** RBFFD | Ecuación de Helmholtz | Parámetro de forma | Efecto polución

#### **AFFILIATION**

°Instituto de Matematicas, Universidad de Antioquia Calle 67 53-108, Medellin, Colombia \*email: alejandro.londono@udea.edu.co



## **1** INTRODUCTION

The Helmholtz equation is an elliptic partial differential equation that represents time-independent solutions of the wave equation. This equation models a wide variety of physical phenomena. These include among others, acoustic wave scattering, time harmonic acoustic, electromagnetic elds, water wave propagation, membrane vibration and radar scattering. All these applications make getting accurate numerical solution for the Helmholtz equation the object of a large number of investigations and methods such as finite differences spectral elements, finite elements method and boundary elements.

Intended for of seismic modeling and inversion, this paper deals with the problem of acoustic and constant density wave propagation on a rectangular domain  $\Omega$  by solving the Helmholtz equation.

$$-\Delta u(\mathbf{x}) - \omega^2 c(\mathbf{x})^{-2} u(\mathbf{x}) = f(\mathbf{x})$$
<sup>(1)</sup>

where  $\omega$  is angular frequency, c(x) > 0 is a smooth function that represents the propagation speed of the wave and f(x) is a compactly supported distribution in  $\omega$  that represents the source distribution. We consider boundary conditions (Dirichlet, Neumann and absorbing boundary conditions [5]) on the border  $\delta\Omega$  and perfectly matched layers(PML) [1],[16] on an artificial extended domain a  $\Omega_\delta \supset \Omega.$ 

When solving (1) at high frequencies by standard numerical methods the wavelength of the numerical solution is different than the real one  $\lambda(x)=2\pi/k(x)$ , where  $k(x)=\omega c(x)^{-1}$  is the so-called wavenumber. This discrepancy of wavelengths is known as pollution effect[2],[17]. To obtain a better solution, it is necessary to oversample the solution by increasing the number of points per wavelength  $N_g$ , which has as drawback more computational cost, time and memory. For these and other facts the design of fast and stable solvers for Helmholtz equation has become a significant computational challenge.

Seismic information of good quality about the earth's subsurface structures is the rationale for geophysical applications. The oil and gas industry uses computational intensive algorithms to provide an image of the subsurface. Information is obtained by sending elastic energy into the subsurface and recording the signal required for a seismic wave to be reflected back to the surface from the Earth interfaces that may have different physical properties. In the petroleum industry, accurate seismic information for such a structure can help in determining potential oil and gas reservoirs in subsurface layers. The seismic wave is usually generated by shots of known frequencies, placed on the earth's surface, and the returning wave is recorded by instruments also placed along the earth's surface.

Helmholtz equation solutions are essential in frequency domain seismic imaging methods such us full waveform inversion (FWI), which requires to solve this equation several times [34]. The traditional Finite Element Method (FEM) or standard Finite Differences (FD) produce sparse matrices but for a few degrees of freedom, there may be a strong pollution effect for large values of the wavenumber. [3],[4],[38] obtain solvers that mitigate the pollution effect in the high frequency regime. Recently, [13] developed a solver by using numerical micro-local analysis and ray-based finite elements method to achieve accurate local finite element space. Furthermore, global radial basis functions are more accurate but produce full matrices that become unstable.

Recently, [25] proposed a RBF-FD solver for frequency domain wave propagation and suggested that the approximated solution does not disperse at relatively high frequencies. In [8] There is an up to date view of the theory and applications of Radial Basis Functions and RBF-FD methods in the numerical solution of partial differential equations with examples in Geosciences and seismic problems.

This paper develops and applies a RBF-FD method on hexagonal grids that solves the Helmholtz equation for a wide range of values of the wavenumber k, focusing in obtaining accurate local approximations of the partial derivative operators by using a few degrees of freedom. Direct solvers may be applied to the banded matrices obtained; in particular we use UMFPACK.

## **2. THEORICAL FRAMEWORK**

#### **RBF INTERPOLATION**

interpolation with Radial Basis Functions the objective is to reconstruct a d-variate function u defined on a bounded domain  $\Omega \subset \mathbb{R}^d$  from the values  $u(x_k)$  of u on a finite set of N scattered points  $X = \{x_1, x_2, \dots, x_N\} \subset \Omega \subset \mathbb{R}^d$ . A radial basis function with shape parameter  $\varepsilon$  is defined as  $\Phi_{\varepsilon}: \mathbb{R}^d \times \mathbb{R}^d \to \mathbb{R}$  [6],[31], [34] such that  $\Phi_{\varepsilon}(x_i y) = \Phi(\varepsilon | |x-y||)$ , where  $\Phi: [0, \infty) \to \mathbb{R}$  is an univariate function. A sufficiently smooth function  $u: \mathbb{R}^d \to \mathbb{R}$  can be approximated by the interpolant.

$$\mathcal{P}_{X,\varepsilon}u(\mathbf{x}) = \sum_{j=1}^{N} \alpha_j \Phi_{\varepsilon}(\mathbf{x}, \mathbf{x}_j).$$
<sup>(2)</sup>

By forcing the condition  $P_{X,\varepsilon}u(x_k) = u(x_k)$  for k = 1,...,N, the weights  $\alpha_j$  can be determined by solving the linear system.

$$u(\mathbf{x}_k) = \sum_{j=1}^N \alpha_j \Phi_{\varepsilon}(\mathbf{x}_k, \mathbf{x}_j), \text{ with } k = 1, \dots, N, \quad (3)$$

provided that the interpolation matrix  $\Phi_{x,\varepsilon} = \Phi_{\varepsilon}(x_k, x_k))_{1 \le k, j \le N}$  is nonsingular and hence the interpolant  $P_{x,\varepsilon}u$  with

$$\begin{pmatrix} \alpha_1 \\ \vdots \\ \alpha_N \end{pmatrix} = (\Phi_{X,\varepsilon})^{-1} u|_X.$$
 (4)

In particular, applying the Gaussian function  $\Phi(r)=e^{-r^2}$ , the interpolation matrix  $\Phi_{X,\varepsilon}$  is positive definite. Some important results about the convenience and properties of interpolation with this function can be found in [6].

The order of the approximation  $u(x) \approx P_{X_{\varepsilon}}u(x)$  can be measured by the fill distance of X in  $\Omega$  defined by

$$h_{X,\Omega} = \sup_{\mathbf{x}\in\Omega} \min_{1\le j\le N} \|\mathbf{x} - \mathbf{x}_j\|$$
(5)

and the stability [34] by the separation distance of *X* defined by

$$q_X = \frac{1}{2} \min_{1 \le j \ne k \le N} \|\mathbf{x}_j - \mathbf{x}_k\|.$$
(6)

It is worth noting that RBF interpolation is meshless because it can be applied to scattered data with no dependence on point distribution. This property and the values of the shape parameter  $\varepsilon$  are commonly used to choose a geometry that may improve some performance or relevant aspects of the problem.

#### **GLOBAL COLLOCATION METHOD WITH RBF**

In recent years there has been increasing interest for obtaining approximated solutions of partial differential equations by the collocation method via RBF interpolation, either by unsymmetrical collocation or symmetric collocation [29]. Details for the latter can be consulted in [35]. Next, the unsymmetrical method will be described [19]. Under the RBF interpolation framework, we want to approximate the solution of boundary value problems in the formula

$$\begin{cases} \mathcal{L}u(\mathbf{x}) = f(\mathbf{x}) & \mathbf{x} \in \Omega \\ \mathcal{B}u(\mathbf{x}) = g(\mathbf{x}) & \mathbf{x} \in \partial\Omega, \end{cases}$$
(7)

where  $\mathcal{L}$  and  $\mathcal{B}$  are linear partial differential operators with regular coefficients have a regularity that is good enough. It is assumed that (7) it is a well-posed problem. Let  $X = [x_i]_{i=1}^m \subset \Omega \cup \partial \Omega$  be a set of points conveniently separated as  $[x_i]_{i=1}^m \subset \Omega$  and  $[x_i]_{i=1}^N \subset \partial \Omega$ . If it is supposing that the unique solution of (7) can be approximated by the interpolant  $P_{X_E}$  in (2), the problem may be forced by the equations

$$\begin{cases} \sum_{j=1}^{N} \alpha_j \mathcal{L} \Phi_{\varepsilon}(\mathbf{x}, \mathbf{x}_j) |_{\mathbf{x}=x_k} = f(\mathbf{x}_k) & \mathbf{x}_k \in \Omega\\ \sum_{j=1}^{N} \alpha_j \mathcal{B} \Phi_{\varepsilon}(\mathbf{x}, \mathbf{x}_j) |_{\mathbf{x}=x_k} = g(\mathbf{x}_k) & \mathbf{x}_k \in \partial\Omega, \end{cases}$$
(8)

arising from the linear system

$$\begin{pmatrix} \mathcal{L}\Phi_{\varepsilon,X\cap\Omega} \\ \mathcal{B}\Phi_{\varepsilon,X\cap\partial\Omega} \end{pmatrix} \begin{pmatrix} \alpha_1 \\ \vdots \\ \alpha_N \end{pmatrix} = \begin{pmatrix} f|_{X\cap\Omega} \\ g|_{X\cap\partial\Omega} \end{pmatrix}$$
(9)

and by (4), obtaining the following linear system whose unknowns are the u values at the points of X

$$\begin{pmatrix} \mathcal{L}\Phi_{\varepsilon,X\cap\Omega} \\ \mathcal{B}\Phi_{\varepsilon,X\cap\partial\Omega} \end{pmatrix} (\Phi_{\varepsilon,X})^{-1} u|_X = \begin{pmatrix} f|_{X\cap\Omega} \\ g|_{X\cap\partial\Omega} \end{pmatrix}$$
(10)

To get the approximated solution  $\tilde{u}$  from (10), it is necessary that the collocation matrix  $(\mathcal{L}\phi_{\epsilon,xn\Omega} \mathcal{B}\phi_{\epsilon,Xn\partial\Omega})^t$  be non-singular, which in general is not true. An elaborated counterexample using multiquadrics and Gaussians was presented in [15]. However, under certain conditions, unsymetrical collocation method is feasible from a generalized approach using separated trial and test spaces [23],[24].

#### **DISCRETIZATION BY RBF-FD**

For problems requiring to compute solutions on large domains good resolution, the resultant matrix obtained by the collocation method is dense, huge and ill-conditioned, implying a prohibited computational cost. A variant of RBF collocation method enabling to deal with large domains is the local version [33],[36]. A local interpolation is possible to obtain a sparse matrix that represents the linear partial differential operator. This approach is often called Radial Basis Function-generated Finite Differences method (RBF-FD). The RBF-FD method is the following.

Let  $X = [x_i]_{i=1}^n \subset \Omega \cup \partial \Omega$  be a set of interpolation points. For any  $x_i \in X$  an influence domain  $S_i \subset X$  is created which is formed by the  $n_i$  nearest neighbor interpolation points. That is, consider an  $n_i$ -stencil  $S_i = [x_i]_{i=1}^{n_i} \subset X$ , where  $x_i' \equiv x_i$  and  $\Omega_i = \text{ConvexHull}(S_i)$ . This is how the subsets  $[S_i]_{i=1}^{N}$  of points contained in X. u(x) are formed, which be approximated by RBF interpolation for  $x \in S_i$  as

$$u(\mathbf{x}) \approx \widetilde{u}(\mathbf{x}) = \mathcal{P}_{S_i,\varepsilon_i} u(\mathbf{x}) = \sum_{j=1}^{n_i} \alpha_j^i \Phi_{\varepsilon_i}(\mathbf{x}, \mathbf{x}_j^i).$$
(11)

with  $x \in \Omega_i$ . On the shape parameter  $\varepsilon_i$  is on depending on the location  $x_i$ . This allows to manipulate the shape of the RBF according to known data. Collocating the  $n_i$  points of the stencil  $S_i$ , results in a small linear system given by

$$\Phi_{S_i,\varepsilon_i} \boldsymbol{\alpha}^i = U_i \tag{12}$$

with  $U_i = (\tilde{u}(x_1^i) \ \tilde{u}(x_2^i) \dots \ \tilde{u}(x_{nl}^i))^t \cdot \Phi_{\varepsilon_k,S_l} = (\Phi_{\varepsilon_l}(x_p^i, x_k^i))_{1 \le j,k \le n_l}$  is the local interpolation matrix and  $\alpha^{i=}(\alpha_{1,1}^i, \alpha_{2,\dots,n}^i, \alpha_{n_l})^t$ .

The unknown coefficients  $\alpha^i$  in (12) can be expressed in terms of the function values at the local interpolation points as

$$\alpha^{i} = \Phi_{S_{i},\varepsilon_{i}}^{-1} U_{i}. \tag{13}$$

The inverse matrix  $\Phi_{3_{k}\epsilon_{i}}^{1}$  exists because of the positive definiteness of  $\Phi_{s_{k}\epsilon_{i}}[6]$ . Now, with the aim of obtaining a local discretized version for (7) consider  $x_{i} \in \Omega \cap X$  or  $X_{i} \in \partial \Omega \cap X$ . In both cases a linear partial differential operator must be applied, either  $\mathcal{L}$  or  $\mathcal{B}$ , to equation (11). For  $x_{i} \in \Omega$  it follows

$$\mathcal{LP}_{S_{i},\varepsilon_{i}}u(\mathbf{x}_{i}) = \sum_{j=1}^{n_{i}} \alpha_{j}^{i} \mathcal{L}\Phi_{\varepsilon_{i}}(\mathbf{x}_{i}, \mathbf{x}_{j}^{i})$$
$$= \mathcal{L}\Phi_{S_{i},\varepsilon_{i}}^{1} \alpha^{i}$$
$$= \mathcal{L}\Phi_{S_{i},\varepsilon_{i}}^{1} \Phi_{S_{i},\varepsilon_{i}}^{-1} U_{i}$$
(14)

where

 $\begin{array}{l} \mathcal{L} \ \varphi_{s_{i}\varepsilon_{i}}^{-1} = (\mathcal{L} \varphi_{\varepsilon_{i}}(x_{\nu}^{i} x_{1}^{i}) \dots \mathcal{L} \varphi_{\varepsilon_{i}}(x_{\nu}^{i} x_{n_{i}}^{i})). \\ \text{Similarly, for } X_{i} \in \partial \Omega \cap X \end{array}$ 

$$\mathcal{BP}_{S_i,\varepsilon_i}u(\mathbf{x}_i) = \mathcal{B}\Phi^1_{S_i,\varepsilon_i}\Phi^{-1}_{S_i,\varepsilon_i}U_i.$$
(15)

We denote  $\mathcal{L}_{S_{j\epsilon ij}} = \mathcal{L} \Phi_{S_{j\epsilon i}}^{-1} \Phi_{S_{j\epsilon i}}^{-1} \Phi_{S_{j\epsilon i}}^{-1}$  and  $\mathcal{B}_{S_{j\epsilon i}} = \mathcal{B}_{S_{j\epsilon i}}^{-1} \Phi_{S_{j\epsilon i}}^{-1}$  From (14) and (15) it follows a discretized local version of (7)

$$\mathcal{L}_{S_i,\varepsilon_i} U_i = f(\mathbf{x}_i), \text{ for } \mathbf{x}_i \in \Omega \cap X, \\ \mathcal{B}_{S_i,\varepsilon_i} U_i = g(\mathbf{x}_i), \text{ for } \mathbf{x}_i \in \partial \Omega \cap X. \end{cases}$$

The above system of linear equations can be assembled forming sparse matrix H of size  $N \times N$  where the *i*-th row, associated to  $x_i \in X$ , has at most ni nonzero entries, and the unknown matrix is given by  $U = \tilde{u}(x_1) \tilde{u}(x_2) \cdots \tilde{u}(x_N)$  (recall  $x_i \equiv x_1^i$ ).

An approximated solution  $\tilde{u}(x)$  at all the interpolation points can be obtained by solving HU = F. H can be considered a discretized version of the operator  $\mathcal{L}$  including the boundary conditions involved in (7). The foregoing discussion gives rise to the following definition.

**Definition 1.** Let  $\mathcal{L}: C^k(\Omega) \to C^{k:m}(\Omega)$  be a linear partial differential operator of order m, where  $\Omega \subset \mathbb{R}^d$  is open and bounded. For  $u \in C^k(\Omega)$  and  $x \in \Omega$ consider an n-stencil  $S_x = \{x_i\}_{i=1}^n \subset \Omega \cup \partial \Omega$  of nearest neighbors to x. Here  $x_i \equiv x$ . The Gaussian RBF-FD operator associated to L denoted by  $\mathcal{L}S_{x_E}$  is defined by

$$\mathcal{L}_{S_{\mathbf{x}},\varepsilon}u(\mathbf{x}) = \mathcal{L}\Phi^{1}_{S_{\mathbf{x}},\varepsilon}\Phi^{-1}_{S_{\mathbf{x}},\varepsilon}U_{S_{\mathbf{x}}},$$
(16)

where

$$\mathcal{L}\Phi_{S_{\mathbf{x}},\varepsilon}^{1} = \left(\mathcal{L}\Phi_{\varepsilon}(\mathbf{y},\mathbf{x}_{1})|_{\mathbf{y}=\mathbf{x}_{1}} \mathcal{L}\Phi_{\varepsilon}(\mathbf{y},\mathbf{x}_{2})|_{\mathbf{y}=\mathbf{x}_{1}} \\ \cdots \mathcal{L}\Phi_{\varepsilon}(\mathbf{y},\mathbf{x}_{n})|_{\mathbf{y}=\mathbf{x}_{1}}\right)$$
(17)

$$\Phi_{S_{\mathbf{x},\varepsilon}} = \begin{pmatrix} \Phi_{\varepsilon}(\mathbf{x}_{1},\mathbf{x}_{1}) & \cdots & \Phi_{\varepsilon}(\mathbf{x}_{1},\mathbf{x}_{n}) \\ \vdots & \ddots & \vdots \\ \Phi_{\varepsilon}(\mathbf{x}_{n},\mathbf{x}_{1}) & \cdots & \Phi_{\varepsilon}(\mathbf{x}_{n},\mathbf{x}_{n}) \end{pmatrix}$$
(18)

$$U_{S_{\mathbf{x}}} = \left( \begin{array}{ccc} u(\mathbf{x}_1) & u(\mathbf{x}_2) & \cdots & u(\mathbf{x}_n) \end{array} \right)^t$$
 (19)

and  $\Phi_{\epsilon}(y;x_i) = e^{-\epsilon^2} ||y-x_i||^2$  with  $\epsilon > 0$ . The Gaussian weights matrix of  $\mathcal{L}$  at  $S_x$  is defined by

$$W\mathcal{L}_{S_{\mathbf{x}},\varepsilon} = \mathcal{L}\Phi^{1}_{S_{\mathbf{x}},\varepsilon}\Phi^{-1}_{S_{\mathbf{x}},\varepsilon}.$$
(20)

initially the application of Gaussian RBF-FD implies the solution of a small linear system at every node in order to compute weight matrices, a condition that increases the computational complexity of the problem. There are situations where the non-singular matrix  $\Phi_{S_{p}e}$ is ill-conditioned, principally for small values of  $\varepsilon$ . In these events the RBF is near to the flat limit, however this issue has been overcome by using different methodologies resulting in stable calculations of (20), for example [10]-[12],[22], and recently [26] have used an hybrid kernel that is formed with two weighted terms, one Gaussian and another cubic which is given by ' $\varphi(r) = \alpha e^{-(\varepsilon r)^2} + \beta r^3$ , so this new hybrid kernel now involves three parameters.

Some RBF's do not contain a shape parameter  $\varepsilon$ . One well known example are polyharmonic splines  $\varphi(r) = r^{2m} \log r, m \in \mathbb{N}$  and  $\varphi(r) = r^{2m}; m \notin \mathbb{N}$ . In general a RBF interpolator has the form

$$P_X u(\mathbf{x}) = \sum_{j=1}^N \alpha_j \phi(||\mathbf{x} - \mathbf{x}_j||) + p(\mathbf{x}),$$

where p(x) is a low degree polynomial. The inclusion of p(x), determines the invertibility of the interpolation matrix and this can be explained in terms of the theory of conditionally positive definite functions (for a wider perspective the reader may consult [34],[6]). For example polyharmonic splines include p(x), while the gaussian function does not.

Up to now two kinds of RBF's have been applied in RBF-FD methods (i) infinitely smooth RBF's (which is the approach used in this paper) which contains a shape parameter. In this case is possible to get a high degree of accuracy in the solution although it is necessary to design strategies or algorithms for tuning the best value of the shape parameter and (ii) polyharmonic splines combined with the polynomial term RBF(PHS)- FD for short. This combination is free from a shape parameter but depends on the choice of PHS and polynomial degrees.  $\Phi(r)$  is less in uential but p(x) cannot exceed a certain maximum, which increases with the local support size. Thus, the error estimates of RBF(PHS)-FD will depend on knowing the maximal permissible degree of supplementary polynomials (see Fornberg's book [8] for further information).

In this work, to obtain a relatively cheap solver, we focused on obtaining predictable weights with closed formulas depending on  $\varepsilon$  and h,  $\varepsilon$  will be adaptable according to the minimum local truncation error at position x. So we will be exploring within small stencils of node sets distributed on regular grids. The node density distribution may cause problems as to accuracy of the method. [8] discuss some methods in quasi-uniform node sets. In order to avoid point sets with sharp density differences we apply the algorithm from [30].

# SHAPE PARAMETER FOR SOLUTIONS OF HELMHOLTZ EQUATION

For an homogeneous media the propagation of time-harmonic waves is modeled by Helmholtz equation with wavenumber  $k = \omega c^{-1}$  where *c* is the sound speed and  $\omega$  is the angular frequency. From the Fourier optics standpoint a time-harmonic wave can be seen as a superposition of plane waves and from the stationary phase method the timeharmonic wave field, at distant points, is due to a plane wave component; therefore it makes sense to consider planes waves as a key part of the analysis. Wave planes are a fundamental tool for investigating the behavior of numerical solutions for Helmholtz equation, either to reduce pollution effects or to design higher order discretizations, in this token, there are recent works such as [18] and [13].

Described below the method to obtain adequate shape parameters on general stencils for approximating a linear dierential operator  ${\cal L}$  with smooth coefficients.

$$\mathcal{L} = \sum_{r,s} l_{rs}(\mathbf{x}) \partial_x^r \partial_y^s$$

when it is applied to solutions of Helmholtz equation. Let  $p_{\mathcal{L}}(\xi, \eta) = \sum_{r,s} I_{r,s}(x) \xi \, {}^{r} \eta^{s}$  be the symbol of the linear differential operator  $\mathcal{L}$  [37]<sup>2</sup>. In particular consider plane waves given by  $u(x; K, \xi \theta) = e^{ik(x\cos\theta + y\sin\theta)}$ , therefore *u* satises

$$\mathcal{L}u(\mathbf{x};k,\theta) = p_{\mathcal{L}}(ik\cos\theta, ik\sin\theta)u(\mathbf{x};k,\theta).$$
 (21)

Let  $S = \{x_i\}_{i=1}^n$  be an *n*-stencil with respect to  $x_i = x = (x, y)$ . The local truncation error by the approximation  $\mathcal{L}_{S,c}u(x; k; \theta)$  is given by

$$\mathcal{L}u(\mathbf{x};k,\theta) - \mathcal{L}_{S,\varepsilon}u(\mathbf{x};k,\theta)$$
  
=  $p_{\mathcal{L}}(ik\cos\theta, ik\sin\theta)u(\mathbf{x};k,\theta) - \mathcal{L}_{S,\varepsilon}u(\mathbf{x};k,\theta)$   
=  $u(\mathbf{x};k,\theta)(p_{\mathcal{L}}(ik\cos\theta, ik\sin\theta) - \frac{1}{u(\mathbf{x};k,\theta)}\mathcal{L}\Phi_{S,\varepsilon}^{1}\Phi_{S,\varepsilon}^{-1}U_{S})$   
=  $u(\mathbf{x};k,\theta)(p_{\mathcal{L}}(ik\cos\theta, ik\sin\theta) - \mathcal{L}\Phi_{S}^{1}\Phi_{S}^{-1}U_{S})$ 

where  $U_{s_0} = (u(x_1 - x; k; \theta) \cdots u(x_n - x; k; \theta))^t$ . Note that

$$\begin{aligned} |\mathcal{L}u(\mathbf{x};k,\theta) - \mathcal{L}_{S,\varepsilon}u(\mathbf{x};k,\theta)| \\ &= |p_{\mathcal{L}}(ik\cos\theta, ik\sin\theta) - \mathcal{L}\Phi_{S,\varepsilon}^{1}\Phi_{S,\varepsilon}^{-1}U_{S_{0}}| \end{aligned}$$
(22)

does not depend explicitly of x. We introduce the averaged local truncation error  $\delta \mathcal{L}_{s}(\varepsilon, k)$  over all directions  $\theta$ , by the integral

$$\delta \mathcal{L}_{S}(\varepsilon,k) = \frac{1}{2\pi} \int_{0}^{2\pi} |p_{\mathcal{L}}(ik\cos\theta, ik\sin\theta) - \mathcal{L}\Phi_{S,\varepsilon}^{1}\Phi_{S,\varepsilon}^{-1}U_{S_{0}}|d\theta.$$
(23)

From the last formula we choose the best shape parameter  $\varepsilon_{\mathcal{LS}}(k)$ , by

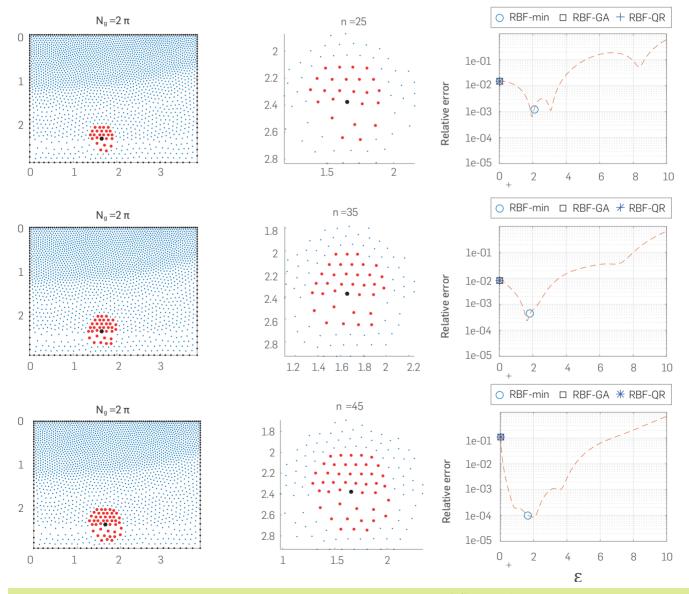
$$\varepsilon_{\mathcal{L},S}(k) = \underset{\varepsilon>0}{\operatorname{arg\,min}} \, \delta\mathcal{L}_S(\varepsilon,k) \tag{24}$$

In particular, if  $\mathcal{L} = \partial_k \partial_{\delta_1}^x$ , then  $p\mathcal{L}(ik\cos\theta; ik\sin\theta) = (ik)^{r*s}\cos^r\theta\sin^s\theta$ . Also, since the symbol of Laplace operator is the polynomial  $p\Delta$  ( $\xi, \eta$ ) =  $\xi^2 + \eta^2$ , then  $p\Delta(ik\cos\theta; ik\sin\theta) = -k^2$ .

For implementations purposes, we approximated the integral (23) by means of Simpson's rule with  $\Delta \theta = 2\pi/16$ . To estimate the minimum argument (24) we used the Matlab function fminbnd.m, which use both Golden-section search and parabolic interpolation. This estimate requires roughly range between ten and twenty evaluations of (23). At each iteration,  $\mathcal{L} \Phi_{\mathcal{S}e}^{1} \Phi_{\mathcal{S}e}^{1}$  is calculated once by Cholesky decomposition. With this procedure we try to reduce dispersion and pollution effects in numerical solutions of Helmholtz equation, calling this algorithm RBF-min.

#### **DISPERSION ANALYSIS**

Consider a solution *u* of Helmholtz equation  $-\Delta u(x)-k^2u(x)=0$ . The classical dispersion analysis applies the relative phase error  $|k-k_f|/k$ 



**Figure 1**. Local truncation error approximating  $\Delta u_i(x_c)$  for several stencil size.  $u_i(x) = \frac{1}{4}H_0^{(1)}(k|x-xs|) + \frac{1}{15}u(x;k,\pi/4)$ , in the figures Xc is the biggest point in black; at  $x_o$  k = 9; 56 and Xs = (3:5; 0:5). Node distribution is based on a region of the smooth Marmousi model.

where  $k_{\rm f}$  is the wavenumber of the numerical solution, which locally satises

$$-\Delta_{S,\varepsilon}u(\mathbf{x}) - k_f^2 u(\mathbf{x}) = 0.$$
(25)

The occurrence of the fictitious wavenumber  $k_f$  is known as numerical dispersion and can be studied by the phase lag  $|k-k_f|$ . However, as the solution of the Helmholtz equation fluctuates considerably for large wavenumbers the situation is even worse because for measuring the properties of a finite difference scheme, considering only the convergence order is not enough. In fact, the accuracy of the numerical solution deteriorates with increasing wavenumber k even when the resolution factor kh, is kept constant. This phenomenon is related to the so-called 'pollution effect' (see, [17]). Therefore our next task is to build an strategy to mitigate this phenomenon.

To obtain an explicit relation between k and  $k_f$  we consider the plane wave

$$u(x, y; k, \theta) = e^{ik((\cos \theta)x + (\sin \theta)y)}$$

with propagation angle  $\theta$  and wave length  $\lambda = 2\pi/k$ . According to (25) we can take  $k_f$  as a function which depends on k,  $\theta$  the stencil S and  $\varepsilon$ , given by its square

$$k_f^2(k,\theta,\varepsilon) = -\frac{\Delta_{S,\varepsilon}u(x,y;k,\theta)}{u(x,y,k,\theta)}.$$
(26)

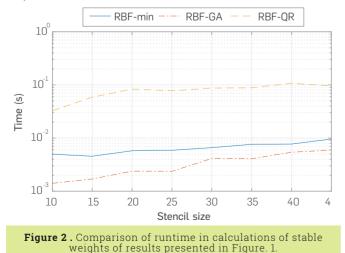
Note that the local truncation error (22) for  $\mathcal{L}=\Delta$  is given just by

$$|\Delta u(\mathbf{x};k,\theta) - \Delta_{S,\varepsilon} u(\mathbf{x};k,\theta)| = |k^2 - k_f^2(k,\theta,\varepsilon)|, (27)$$

The fictitious wavenumber is anisotropic, i.e., it depends on the wave propagation direction  $\theta$ . As was the case in [27], our objective will be to choose the shape parameter  $\varepsilon$  such that the average phase error, over all angles of propagation, is minimum. Hence in minimizing the averaged local truncation error

$$\delta\Delta_{S}(\varepsilon,k) = \frac{1}{2\pi} \int_{0}^{2\pi} |k^{2} - \Delta\Phi_{S,\varepsilon}^{1}\Phi_{S,\varepsilon}^{-1}U_{S_{0}}|d\theta \quad (28)$$

for the variable  $\boldsymbol{\epsilon}$  we can mitigate the pollution effect due to dispersion error



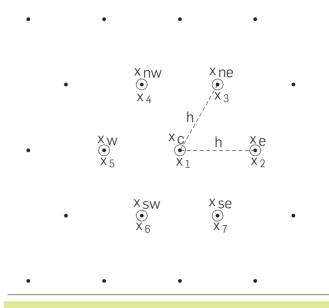


Figure 3.7-stencil on a regular hexagonal grid.

#### SOME CLOSED FORMULAS AND TRUNCATION ERROR

In the context of numerical solutions of the wave equation, numerical stability and isotropy are enhanced with a similar computational work when compared with the standard numerical methods in the simulation of heterogeneous and random media. For example, whereas Von Neumann stability in solution of the 2D acoustic wave equation by explicit finite differences with 5–points needs to satisfy the estimate  $\frac{\Delta t}{\Delta x} \leq c \cdot \frac{1/2}{V_3}$ , with hexagonal 7–stencil, this estimate is improved to  $\frac{\Delta t}{\Delta x} \leq c \cdot \frac{1/2}{V_3}$  [7]. Besides, with hexagonal stencils we have reached similar results to those reported in [3] where the authors have developed an optimal Cartesian 9–stencil for Helmholtz equation with the goal of mitigating the pollution effect.

We start considering the 7-stencil  $H = [x_i]_{i=1}^{r}$ , built on an hexagonal regular grid as in **Figure 3**. For h > 0 and  $x_i = (x; y)$  as central node, the coordinates of the other points in *H* are given by

$$\mathbf{x}_{2} = (x+h,y), \ \mathbf{x}_{3} = (x+h/2, y+\sqrt{3}h/2), \\ \mathbf{x}_{4} = (x-h/2, y+\sqrt{3}h/2), \ \mathbf{x}_{5} = (x-h,y), \\ \mathbf{x}_{6} = (x-h/2, y-\sqrt{3}h/2), \ \mathbf{x}_{7} = (x+h/2, y-\sqrt{3}h/2)$$

Sometimes will be convenient to work with subscripts referred by cardinal directions:

$$\begin{aligned} \mathbf{x}_c &= \mathbf{x}_1, \ \mathbf{x}_e &= \mathbf{x}_2, \ \mathbf{x}_{ne} &= \mathbf{x}_3, \ \mathbf{x}_{nw} &= \mathbf{x}_4 \\ \mathbf{x}_w &= \mathbf{x}_5, \ \mathbf{x}_{sw} &= \mathbf{x}_5, \ \mathbf{x}_{se} &= \mathbf{x}_7 \end{aligned}$$

With the given labeled order in the 7-stencil *H* and with  $p = e^{-(\varepsilon h)^2}$ , the interpolation matrix  $\Phi_{H,\varepsilon}$  is explicitly given by

$$\Phi_{H,\varepsilon} = \begin{pmatrix} 1 & p & p & p & p & p & p \\ p & 1 & p & p^3 & p^4 & p^3 & p \\ p & p & 1 & p & p^3 & p^4 & p^3 \\ p & p^3 & p & 1 & p & p^3 & p^4 \\ p & p^4 & p^3 & p & 1 & p & p^3 \\ p & p^3 & p^4 & p^3 & p & 1 & p \\ p & p & p^3 & p^4 & p^3 & p & 1 \end{pmatrix}.$$
(29)

From this formula, we seek expressions for  $(\partial_x)_{H,\varepsilon}$ ,  $(\partial_y)_{H,\varepsilon}$ ,  $(\partial_x^2)_{H,\varepsilon}$ ,  $(\partial_x^2)_{H,\varepsilon}$ ,  $(\partial_y^2)_{H,\varepsilon}$ ,  $(\partial_y^2)_{H,\varepsilon}$ ,  $(\partial_x)_{H,\varepsilon}$ ,  $(\partial_x)_{H,\varepsilon}$ ,  $(\partial_x)_{H,\varepsilon}$ ,  $(\partial_y^2)_{H,\varepsilon}$ ,  $(\partial_y^$ 

where

$$D(h,\varepsilon) = \frac{h\varepsilon^{2} \left( \coth \left(h^{2}\varepsilon^{2}\right) + 1 \right)}{4 \cosh \left(h^{2}\varepsilon^{2}\right) + 2}$$

$$E(h,\varepsilon) = \frac{3h^{2}\varepsilon^{4}}{2 \cosh \left(h^{2}\varepsilon^{2}\right) + \cosh \left(2h^{2}\varepsilon^{2}\right) - 3} + \varepsilon^{2}$$

$$F_{1}(h,\varepsilon) = \frac{h^{2}\varepsilon^{4}e^{h^{2}\varepsilon^{2}} \left(4 \cosh \left(h^{2}\varepsilon^{2}\right) + 5\right) \operatorname{csch}^{2} \left(\frac{h^{2}\varepsilon^{2}}{2}\right)}{4 \left(5 \cosh \left(h^{2}\varepsilon^{2}\right) + \cosh \left(2h^{2}\varepsilon^{2}\right) + 3\right)}$$

$$G_{1}(h,\varepsilon) = \frac{h^{2}\varepsilon^{4}e^{h^{2}\varepsilon^{2}}}{2 \left(5 \cosh \left(h^{2}\varepsilon^{2}\right) + \cosh \left(2h^{2}\varepsilon^{2}\right) + 3\right)}$$

$$F_{2}(h,\varepsilon) = -\frac{3h^{2}\varepsilon^{4}e^{h^{2}\varepsilon^{2}}\operatorname{csch}^{2} \left(\frac{h^{2}\varepsilon^{2}}{2}\right)}{4 \left(5 \cosh \left(h^{2}\varepsilon^{2}\right) + \cosh \left(2h^{2}\varepsilon^{2}\right) + 3\right)}$$

$$G_{2}(h,\varepsilon) = \frac{3h^{2}\varepsilon^{4} \left(e^{h-\varepsilon^{2}} + 1\right)^{2} \operatorname{csch}^{2} \left(\frac{h^{2}\varepsilon^{2}}{2}\right)}{8 \left(5 \cosh \left(h^{2}\varepsilon^{2}\right) + \cosh \left(2h^{2}\varepsilon^{2}\right) + 3\right)}$$

$$F_{3}(h,\varepsilon) = \frac{h^{2}\varepsilon^{4}e^{h^{2}\varepsilon^{2}}\operatorname{csch}^{2} \left(\frac{h^{2}\varepsilon^{2}}{2}\right)}{4 \cosh \left(h^{2}\varepsilon^{2}\right) + 2}$$

Also, we note that  $F_1 + F_2 = G_1 + G_2$  and denote it by  $F = F_1 + F_2 = G_1 + G_2$ , explicitly

$$F(h,\varepsilon) = \frac{h^2 \varepsilon^4 e^{h^2 \varepsilon^2} \operatorname{csch}^2\left(\frac{h^2 \varepsilon^2}{2}\right)}{2 \cosh\left(h^2 \varepsilon^2\right) + 4}$$
(32)

Formulas for partial derivative operators of second order. By simple substitutions with the group of equations (31) within of equations in (30) we have explicit formulas for approximating partial derivative operators of second order

$$(\partial_{x})_{H,\varepsilon}u(\mathbf{x_{c}}) = D(h,\varepsilon)(2u_{2}+u_{3}-u_{4}-2u_{5}-u_{6}+u_{7})$$

$$(\partial_{y})_{H,\varepsilon}u(\mathbf{x_{c}}) = \sqrt{3}D(h,\varepsilon)(u_{3}+u_{4}-u_{6}-u_{7})$$

$$(\partial_{x}^{2})_{H,\varepsilon}u(\mathbf{x_{c}}) = -2E(h,\varepsilon)u_{1}+F_{1}(h,\varepsilon)(u_{2}+u_{5}) + G_{1}(h,\varepsilon)(u_{3}+u_{4}+u_{6}+u_{7})$$

$$(\partial_{y}^{2})_{H,\varepsilon}u(\mathbf{x_{c}}) = -2E(h,\varepsilon)u_{1}+F_{2}(h,\varepsilon)(u_{2}+u_{5}) + G_{2}(h,\varepsilon)(u_{3}+u_{4}+u_{6}+u_{7})$$

$$(\partial_{xy})_{H,\varepsilon}u(\mathbf{x_{c}}) = \frac{\sqrt{3}}{2}F_{3}(h,\varepsilon)(u_{3}-u_{4}+u_{6}-u_{7})$$

$$\Delta_{H,\varepsilon}u(\mathbf{x_{c}}) = -4E(h,\varepsilon)u_{1}+F(h,\varepsilon)(u_{2}+u_{3}+u_{4}+u_{5}+u_{6}+u_{7})$$

As it is usual in finite differences schemes, once an approximation is proposed, the first step is to investigate its local truncation error. The standard methodology consists in making comparisons with Taylor polynomials. The next lemma will be useful in that sense.

**Lemma 1.** For  $0 < \varepsilon h < \sqrt{\frac{\pi}{2}}$  the following approximations of weights functions in (31) hold

$$D(h,\varepsilon) = \frac{1}{6h} + \frac{h\varepsilon^{2}}{6} + \mathcal{O}(\varepsilon^{3}h^{4})$$

$$E(h,\varepsilon) = \frac{1}{h^{2}} + \varepsilon^{2} - \frac{h^{2}\varepsilon^{4}}{4} + \mathcal{O}(\varepsilon^{6}h^{4})$$

$$F_{1}(h,\varepsilon) = \frac{1}{h^{2}} + \varepsilon^{2} + \frac{5h^{2}\varepsilon^{4}}{36} + \mathcal{O}(\varepsilon^{6}h^{4})$$

$$G_{1}(h,\varepsilon) = \frac{h^{2}\varepsilon^{4}}{18} + \frac{h^{4}\varepsilon^{6}}{18} + \mathcal{O}(\varepsilon^{8}h^{6})$$

$$F_{2}(h,\varepsilon) = -\frac{1}{3h^{2}} - \frac{\varepsilon^{2}}{3} + \frac{h^{2}\varepsilon^{4}}{36} + \mathcal{O}(\varepsilon^{6}h^{4})$$

$$G_{2}(h,\varepsilon) = \frac{2}{3h^{2}} + \frac{2\varepsilon^{2}}{3} + \frac{h^{2}\varepsilon^{4}}{9} + \mathcal{O}(\varepsilon^{6}h^{4})$$

$$F_{3}(h,\varepsilon) = \frac{2}{3h^{2}} + \frac{2\varepsilon^{2}}{3} + \frac{h^{2}\varepsilon^{4}}{6} + \mathcal{O}(\varepsilon^{6}h^{4})$$

$$F(h,\varepsilon) = \frac{2}{3h^{2}} + \frac{2\varepsilon^{2}}{3} + \frac{h^{2}\varepsilon^{4}}{6} + \mathcal{O}(\varepsilon^{6}h^{4}).$$

One way to test the quality of the method is to point out the relationship with other finite difference schemes. It is not difficult to show that the standard finite difference method is the limit of the RBF-FD scheme, it occurs when the width of the Gaussian RBF, controlled by the shape parameter  $\varepsilon$ , tends to be at.

**Remark 1.** By examining the limit situation  $\varepsilon \rightarrow 0^+$  with  $\Delta_{H,\varepsilon}u(Xc)$  in (33) via approxi-mations in (34) we recover the formula

$$\Delta_{H,0^+} u(\mathbf{x}_c) = -\frac{4}{h^2} u_c + \frac{2}{3h^2}$$
(35)

 $(u_{\mathbf{e}} + u_{\mathbf{ne}} + u_{\mathbf{nw}} + u_{\mathbf{w}} + u_{\mathbf{sw}} + u_{\mathbf{se}}).$ 

given in [7], corresponding to the standard finite difference formulation derived by Taylor expansions.

Local truncation errors. To estimate the local truncation error of the approximations  $\mathcal{L}_{H,\varepsilon}u(x) \approx \mathcal{L}u(x)$  it suffices to expand  $u(x_i)$ , for i = 2, 3; ..., 7, in Taylor polynomials around  $x_1 = x_c$ , and then substituting them in (33), obtaining

$$\begin{aligned} &(\partial_x)_{H,\varepsilon} u(\mathbf{x}_1) \\ &= 6hD(h,\varepsilon)\partial_x u(\mathbf{x}_1) + \frac{1}{4}D(h,\varepsilon) \left(3h^3(\partial_x^3 u(\mathbf{x}_1) + \partial_x \partial_y^2 u(\mathbf{x}_1)) + \mathcal{O}(h^4)\right) \\ &(\partial_y)_{H,\varepsilon} u(\mathbf{x}_1) \\ &= 6hD(h,\varepsilon)\partial_y u(\mathbf{x}_1) + \frac{1}{4}D(h,\varepsilon) \left(3h^3(\partial_y^3 u(\mathbf{x}_1) + \partial_y \partial_x^2 u(\mathbf{x}_1)) + \mathcal{O}(h^4)\right) \\ &(\partial_x^2)_{H,\varepsilon} u(\mathbf{x}_1) \\ &= \frac{h^2}{2}(2F_1(h,\varepsilon) + G_1(h,\varepsilon))\partial_x^2 u(\mathbf{x}_1) + (2F_1(h,\varepsilon)) \end{aligned}$$

$$= \frac{1}{2} (2F_1(h,\varepsilon) + G_1(h,\varepsilon)) \partial_x^2 u(\mathbf{x}_1) + (2F_1(h,\varepsilon)) + 4G_1(h,\varepsilon) - 2E(h,\varepsilon)) u(\mathbf{x}_1) + \frac{3}{2} h^2 G_1(h,\varepsilon) \partial_y^2 u(\mathbf{x}_1) + F(h,\varepsilon) \mathcal{O}(h^4)$$
(36)

$$\begin{aligned} (\partial_y^2)_{H,\varepsilon} u(\mathbf{x}_1) \\ &= \frac{h^2}{2} (2F_2(h,\varepsilon) + G_2(h,\varepsilon)) \partial_x^2 u(\mathbf{x}_1) + (2F_2(h,\varepsilon)) \\ &+ 4G_2(h,\varepsilon) - 2E(h,\varepsilon)) u(\mathbf{x}_1) \\ &+ \frac{3}{2} h^2 G_2(h,\varepsilon) \partial_y^2 u(\mathbf{x}_1) + F(h,\varepsilon) \mathcal{O}(h^4) \\ (\partial_{xy})_{H,\varepsilon} u(\mathbf{x}_1) \\ &= \frac{3}{2} h^2 F_3(h,\varepsilon) \partial_{xy} u(\mathbf{x}_1) + F_3(h,\varepsilon) \mathcal{O}(h^4) \\ &\Delta_{H,\varepsilon} u(\mathbf{x}_1) = \frac{3}{2} h^2 F(h,\varepsilon) \Delta u(\mathbf{x}_1) + (6F(h,\varepsilon)) \\ &- 4E(h,\varepsilon)) u(\mathbf{x}_1) + 2F(h,\varepsilon) \mathcal{O}(h^4) \end{aligned}$$

From Taylor polynomials in (34) the following error can be obtained

$$\begin{aligned} |6hD(h,\varepsilon) - 1| &\leq h^{2}\varepsilon^{2} \\ \left|\frac{h^{2}}{2}(2F_{1}(h,\varepsilon) + G_{1}(h,\varepsilon)) - 1\right| &\leq h^{2}\varepsilon^{2} + \frac{h^{4}\varepsilon^{4}}{6} \\ |2F_{1}(h,\varepsilon) + 4G_{1}(h,\varepsilon) - 2E(h,\varepsilon)| &\leq h^{2}\varepsilon^{4} \\ \left|\frac{3}{2}h^{2}G_{1}(h,\varepsilon)\right| &\leq \frac{h^{4}\varepsilon^{4}}{12} + \frac{h^{6}\varepsilon^{6}}{12} \\ \left|\frac{h^{2}}{2}(2F_{2}(h,\varepsilon) + G_{2}(h,\varepsilon))\right| &\leq \frac{h^{4}\varepsilon^{2}}{12} + \frac{h^{6}\varepsilon^{6}}{12} \\ \left|\frac{3}{2}h^{2}G_{2}(h,\varepsilon) - 1\right| &\leq h^{2}\varepsilon^{2} + \frac{h^{4}\varepsilon^{4}}{6} \\ |2F_{2}(h,\varepsilon) + 4G_{2}(h,\varepsilon) - 2E(h,\varepsilon)| &\leq h^{2}\varepsilon^{4} \\ \left|\frac{3}{2}h^{2}F(h,\varepsilon) - 1\right| &\leq h^{2}\varepsilon^{2} + \frac{h^{4}\varepsilon^{4}}{4} \\ |6F(h,\varepsilon) - 4E(h,\varepsilon)| &\leq 2h^{2}\varepsilon^{4} \\ \left|\frac{3}{2}h^{2}F_{3}(h,\varepsilon) - 1\right| &\leq h^{2}\varepsilon^{2} + \frac{h^{4}\varepsilon^{4}}{12} \end{aligned}$$

from which the error estimates can be obtained for (33) with their respective approximation errors

$$\begin{split} |(\partial_{x})_{H,\varepsilon}u(\mathbf{x}) - \partial_{x}u(\mathbf{x})| &\leq h^{2}\varepsilon^{2}|\partial_{x}u(\mathbf{x})| + \mathcal{O}(\varepsilon^{2}h^{2}) \\ |(\partial_{y})_{H,\varepsilon}u(\mathbf{x}) - \partial_{y}u(\mathbf{x})| &\leq h^{2}\varepsilon^{2}|\partial_{y}u(\mathbf{x})| + \mathcal{O}(\varepsilon^{2}h^{2}) \\ |(\partial_{x}^{2})_{H,\varepsilon}u(\mathbf{x}) - \partial_{x}^{2}u(\mathbf{x})| &\leq \left(h^{2}\varepsilon^{2} + \frac{h^{4}\varepsilon^{4}}{6}\right)|\partial_{x}^{2}u(\mathbf{x})| \\ &\quad + \left(\frac{h^{4}\varepsilon^{4}}{12} + \frac{h^{6}\varepsilon^{6}}{12}\right)|\partial_{y}^{2}u(\mathbf{x})| \\ &\quad + h^{2}\varepsilon^{4}|u(\mathbf{x})| + \mathcal{O}(\varepsilon^{2}h^{2}) \\ |(\partial_{y}^{2})_{H,\varepsilon}u(\mathbf{x}) - \partial_{y}^{2}u(\mathbf{x})| &\leq \left(\frac{h^{4}\varepsilon^{4}}{12} + \frac{h^{6}\varepsilon^{6}}{12}\right)|\partial_{x}^{2}u(\mathbf{x})| \quad (\mathbf{38}) \\ &\quad + \left(h^{2}\varepsilon^{2} + \frac{h^{\frac{4}{\varepsilon}}}{6}\right)|\partial_{y}^{2}u(\mathbf{x})| \\ &\quad + h^{2}\varepsilon^{4}|u(\mathbf{x})| + \mathcal{O}(\varepsilon^{2}h^{2}) \\ |(\partial_{xy})_{H,\varepsilon}u(\mathbf{x}) - \partial_{xy}u(\mathbf{x})| &\leq h^{2}\varepsilon^{2}|\partial_{xy}u(\mathbf{x})| + \mathcal{O}(\varepsilon^{2}h^{2}) \\ |(\partial_{H,\varepsilon}u(\mathbf{x}) - \Delta u(\mathbf{x})| &\leq \left(h^{2}\varepsilon^{2} + \frac{h^{4}\varepsilon^{4}}{4}\right)|\Delta u(\mathbf{x})| \\ &\quad + 2h^{2}\varepsilon^{4}|u(\mathbf{x})| + \mathcal{O}(\varepsilon^{2}h^{2}). \end{split}$$

The above treatment and discussion can be summarized in the next result

**Theorem 1.** Let *u* be *a* function in  $C^4(\Omega)$  with  $\Omega \subset \mathbb{R}^2$  a convex open set. Then for  $x \in \Omega$ , h > 0 small enough such that  $H \subset \Omega$  and  $\varepsilon > 0$  such that  $0 < \varepsilon h < \sqrt{\frac{\pi}{2}}$ ; the following assertions about the Gaussian RBF-FD approximations hold:

 $\begin{array}{ll} \bullet & (\partial_x)_{H,\varepsilon} u(x) = \partial_x u(x) + \mathcal{O} \ (\varepsilon^2 h^2) \\ \bullet & (\partial_y)_{H,\varepsilon} u(x) = \partial_y u(x) + \mathcal{O} \ (\varepsilon^2 h^2) \\ \bullet & (\partial_x^2)_{H,\varepsilon} u(x) = \partial_x^2 u(x) + \mathcal{O} \ (\varepsilon^2 h^2) \\ \bullet & (\partial_y^2)_{H,\varepsilon} u(x) = \partial_y^2 u(x) + \mathcal{O} \ (\varepsilon^2 h^2) \\ \bullet & (\partial_{xy})_{H,\varepsilon} u(x) = \Delta_x u(x) + \mathcal{O} \ (\varepsilon^2 h^2) \\ \bullet & \Delta_{H,\varepsilon} u(x) = \Delta^2 u(x) + \mathcal{O} \ (\varepsilon^2 h^2) \end{array}$ 

In **Figure 4** it is depicted a comparison of RBF-FD with 5–stencils and 7–stencils against its respective standard finite differences approximations. To conduct this test we chose the Hankel function of the rst kind  $u(x)=H_0^{(1)}(k|x|)$  with k = 100 and h = 0.01. We can observe the behaviour of the relative truncation error

$$\frac{|\partial_x H_0^{(1)}(k|\mathbf{x}_1|) - (\partial_x)_{H \,\varepsilon} H_0^{(1)}(k|\mathbf{x}_1|)|}{|\partial_x H_0^{(1)}(k|\mathbf{x}_1|)|} \text{ (figure 4-left) (39)}$$

and

$$\frac{|\Delta H_0^{(1)}(k|\mathbf{x}_1|) - \Delta_{H,\varepsilon} H_0^{(1)}(k|\mathbf{x}_1|)|}{|\Delta H_0^{(1)}(k|\mathbf{x}_1|)|} \text{ (figure 4-right). (40)}$$

It is worth to point out that in both situations there exist values of  $\epsilon$  where the error reaches a minimum value. Furthermore it confirms that the limit situation  $\epsilon \rightarrow$  0<sup>+</sup>, RBF-FD coincides with FD.

#### **OPTIMAL SHAPE PARAMETER ON HEXAGONAL STENCIL**

We derive closed formulas to rapidly obtain the optimal shape parameter in the context of hexagonal stencils.

Remark 2. Let  $\boldsymbol{u}$  be a solution of the Helmholtz equation. By theorem 1.

$$\Delta_{H,\varepsilon}u(x,y) = \Delta u(x,y) + \mathcal{O}(\varepsilon^2 h^2)$$

for any  $\varepsilon>0$  on a hexagonal stencil H of size h centered at (x;y), we have that for small enough h

$$\Delta_{H,\varepsilon}u(x,y) + k^2u(x,y) = \mathcal{O}(\varepsilon^2 h^2)$$
(41)

In this case the ctitious wavenumber  $k_f$  depends also of the stencil size h, so  $k_f = k_f (k; \theta; h; \varepsilon)$  is such that

$$\Delta_{H,\varepsilon}u(x,y) + k_f^2u(x,y) = 0.$$
(42)

As we have already seen,  $k_f$  is related to the truncation error of the Laplace operator applied in u by  $(k^2 - k_f^2)u(x, y) = O(\epsilon^2 h^2)$ .

In this case, thanks to the symmetry of the hexagonal stencil, it is possible to nd an explicit relation between k and kf . From (33) in (26) we obtain the formula

$$k_f^2(k,\theta,h,\varepsilon) = 4E(h,\varepsilon) - 2F(h,\varepsilon)T(k,\theta,h)$$
 (43)

where

$$T(k,\theta,h) = \cos(hk\cos\theta) + 2\cos\left(\frac{1}{2}hk\cos\theta\right)$$
  
$$\cos\left(\frac{\sqrt{3}}{2}hk\sin\theta\right).$$
 (44)

In Figure 5 (left) there is a plot of the residual  $|\delta k(k,\theta,h,\varepsilon)| = |k^2 - k_f^2(k,\theta,h,\varepsilon)| = |k^2 - k_f^2(k,\theta,h,\varepsilon)|$  where it can appreciated that for some values of  $\varepsilon$  it is possible to have good accuracy even in scaling *h* for keeping hk = 1 or equivalently  $h = \frac{\lambda}{2\pi}$ , that is, by fixing  $2\pi \approx 6$  points per wavelength. Next lemma will allow us to introduce a simplified variant of (43) which does not depend of  $\theta$ , by just taking the mean value over  $[0,2\pi]$ .

Lemma 2. For any  $(h,k) \in \mathbb{R}^2$ , T as in (44) satisfs the equality

$$\int_0^{2\pi} T(k,\theta,h)d\theta = 6\pi J_0(hk),$$
(45)

where  $J_{a}$  is the Bessel function of first kind.

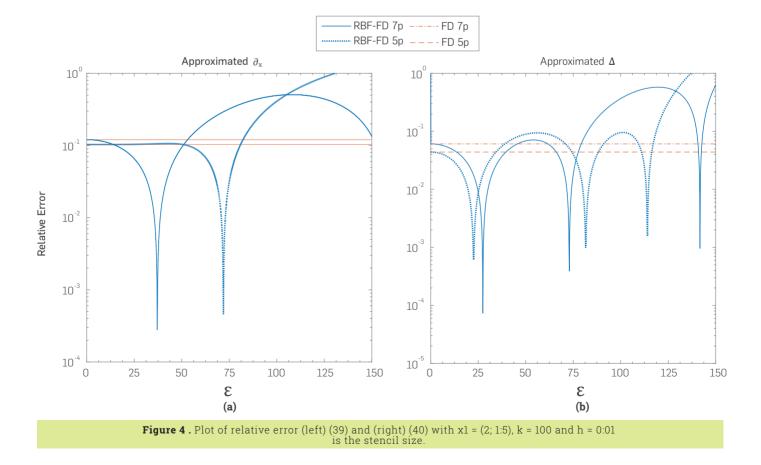
Proof. First, note that  $T(k,\theta + \frac{\pi}{3},h) = T(k,\theta,h)$ , thus T has period  $\frac{\pi}{3}$  in the second argument and

$$\int_0^{2\pi} T(k,\theta,h)d\theta = 6 \int_0^{\pi/3} T(k,\theta,h)d\theta, \qquad (46)$$

then, considering the identity

$$2\cos\left(\frac{1}{2}hk\cos(\theta)\right)\cos\left(\frac{\sqrt{3}}{2}hk\sin(\theta)\right)$$
  
=  $\cos\left(hk\cos\left(\theta + \frac{\pi}{3}\right)\right) + \cos\left(hk\cos\left(\theta - \frac{\pi}{3}\right)\right)$  (47)

and by a few changes of variable we can finally obtain



$$\begin{aligned} &\int_{0}^{\pi/3} T(k,\theta,h) d\theta \\ &= \int_{0}^{\pi/3} \cos(hk\cos\theta) d\theta + \int_{0}^{\pi/3} \cos\left(hk\cos\left(\theta + \frac{\pi}{3}\right)\right) d\theta \\ &+ \int_{0}^{\pi/3} \cos\left(hk\cos\left(\theta - \frac{\pi}{3}\right)\right) d\theta \\ &= \int_{0}^{\pi/3} \cos(hk\cos\theta) d\theta + \int_{\pi/3}^{2\pi/3} \cos(hk\cos\theta) d\theta \\ &+ \int_{2\pi/3}^{\pi} \cos(hk\cos\theta) d\theta \\ &= \int_{0}^{\pi} \cos(hk\cos\theta) d\theta \\ &= \int_{0}^{\pi} \cos(hk\cos\theta) d\theta \\ &= \pi J_{0}(hk) \end{aligned}$$

where the last equality is a well known integral representation of the Bessel functions. We define  $\widetilde{k}_{\rm f}$  such that

$$\widetilde{k}_{f}^{2}(k,h,\varepsilon) = \frac{1}{2\pi} \int_{0}^{2\pi} k_{f}^{2}(k,\theta,h,\varepsilon) d\theta$$
(48)

explicitly results in the formula

$$\widetilde{k}_{f}^{2}(k,h,\varepsilon) = 4E(h,\varepsilon) - 6F(h,\varepsilon)J_{0}(hk), \qquad (49)$$

and we also introduced a new residual function  $\delta \tilde{K}(k,h,\varepsilon)$ . We will see later what is the error between  $k_f(k,h,\varepsilon)$  and  $\tilde{K}_f(k,h,\varepsilon)$  and  $\tilde{K}_f(k,h,\varepsilon)$  produced by "neglecting"  $\theta$  in taking the mean value (48).

**Lemma 3.** Let h and k be positive real numbers with small enough h, then

$$\lim_{\varepsilon \to 0^+} \delta \widetilde{k}(k, h, \varepsilon)$$

$$= \frac{h^2 k^2 + 4J_0(hk) - 4}{h^2} > 0 \text{ and } \lim_{\varepsilon \to \infty} \delta \widetilde{k}(k, h, \varepsilon) = -\infty$$
(50)

Proof. By supposing  $0 < \epsilon h < \sqrt{\frac{\pi}{2}}$ , using the expansions in (34) at (49) and taking  $\epsilon \rightarrow 0^+$ , then  $\lim_{\epsilon \rightarrow 0^+} \delta \widetilde{k}(k,h,\epsilon) = \frac{h^2k^2 + 4J0(hk) - 4}{h^2}$ . From the Taylor series of  $J_0$  it is true that

$$J_0(hk) > 1 - \frac{h^2 k^2}{4},$$

and the inequality is thus obtained. For the second limit note that from the explicit formula for (49) given by

$$\begin{split} &\delta \widetilde{k}(k,h,\varepsilon) = k^2 - 4\varepsilon^2 \left( \frac{3h^2 \varepsilon^2}{2\cosh\left(h^2 \varepsilon^2\right) + \cosh\left(2h^2 \varepsilon^2\right) - 3} + 1 \right) \\ &+ 3\varepsilon^2 \frac{h^2 \varepsilon^2 e^{h^2 \varepsilon^2} \operatorname{csch}^2\left(\frac{h^2 \varepsilon^2}{2}\right)}{\cosh\left(h^2 \varepsilon^2\right) + 2} J_0(hk), \end{split}$$

we can see that

$$\begin{split} \delta \widetilde{k}(0,1,\varepsilon) &= -4\varepsilon^2 \left( \frac{3\varepsilon^2}{2\cosh(\varepsilon^2) + \cosh(2\varepsilon^2) - 3} + 1 \right) \\ &+ 3\varepsilon^2 \frac{\varepsilon^2 e^{\varepsilon^2} \operatorname{csch}^2 \left(\frac{\varepsilon^2}{2}\right)}{\cosh(\varepsilon^2) + 2} \\ &= -4\varepsilon^2 + 6\varepsilon^4 \frac{\operatorname{coth} \left(\frac{\varepsilon^2}{2}\right) + 1}{\cosh(\varepsilon^2) + 2} \end{split}$$

In addition, can be observed above that the behavior of  $\delta \tilde{k}(k,h,\varepsilon)$  and  $\delta \tilde{k}(0,1,\varepsilon)$  when  $\varepsilon \to \infty$  are the same, hence

$$\begin{split} \lim_{\varepsilon \to \infty} \delta \widetilde{k}(k, h, \varepsilon) &= \lim_{\varepsilon \to \infty} \delta \widetilde{k}(0, 1, \varepsilon) \\ &= \lim_{\varepsilon \to \infty} \left( -4\varepsilon^2 + 6\varepsilon^4 \frac{\coth\left(\frac{\varepsilon^2}{2}\right) + 1}{\cosh\left(\varepsilon^2\right) + 2} \right) \\ &= -\infty. \end{split}$$

The above lemma and the continuity of the function  $\delta \tilde{k}(k,h,\varepsilon)$  allow us to guarantee that under the condition h > 0 and h > 0 there is some value  $\varepsilon > 0$  such that  $\delta \tilde{k}(k,h,\varepsilon)=0$ . As we have pointed out,  $\delta k$ measures the dispersion of the approximation, as well  $\delta k$ , thus, given  $h_o$  and  $k_o$ , a good shape parameter would be  $\varepsilon_o(k_o,h_o)$  such that  $\delta k(k_o,h_o,\varepsilon_o(k_o,h_o))=0$ . Our strategy to reduce numerical dispersion is to find the minimum root of the equation  $\delta \tilde{k}(k_o,h_o,\varepsilon)=0$ . For our implementations we will be more restrictive by considering  $\frac{1}{4} \le hk$  $\le 1$ , with this we are choosing h so that for sampling a wavelength  $\lambda = \frac{2\pi}{k}$  it means to use between  $2\pi$  to  $8\pi$  points, that is approximately between 6 and 25 points per wavelength. Thus, in our discretization for the Helmholtz equation on hexagonal stencils we will use the shape parameter

$$\varepsilon_{op}(k,h) = \min\{\varepsilon > 0 : \delta k(k,h,\varepsilon) = 0\}.$$
 (51)

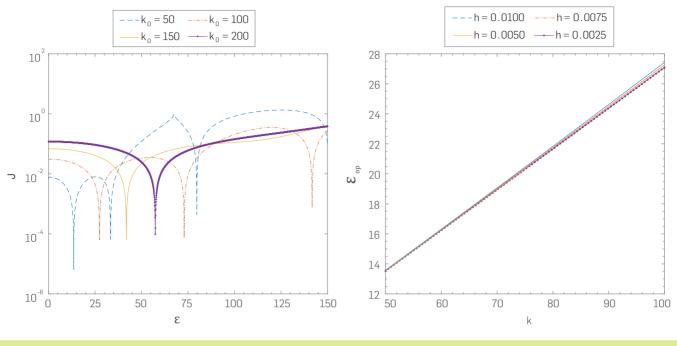
**Figure 5** (Right) shows the results of a test focused on behavior of the optimum shape parameter  $\varepsilon_{op}$ . A linear dependence of the wavenumber k is observed; at least in the applied test.

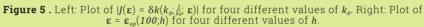
#### ENHANCED LOCAL TRUNCATION ERROR

This section presents the results of a couple of tests intended to show the order of the local truncation error  $|\Delta_{H,\epsilon_{op}}u(x)-\Delta u(x)|$ and  $|(\partial_x)_{H,\epsilon_{op}}u(x)-(\partial_x)u(x)|$ . In **Figure 6** shows the results of testing the approximations by comparing with standard FD for  $u(x)=e^{i\mathbf{k}\cdot \mathbf{x}}$ . Whereas in **Figure 7** the same is done with  $u(x)=\frac{i}{4}H_0^{(1)}(k|\mathbf{x}|)$ . Both functions are solutions of the Helmholtz equation, it can be noted that in these cases the optimum shape parameter oers an error of the order  $\mathcal{O}(h^4)$ .

#### **RELATIVE PHASE ERROR - POLLUTION EFFECT**

As is well known through the vast literature about numerical solutions of Helmholtz equation, k is a key parameter. For large values of k, the solution u is highly fluctuating. To approximate the Helmholtz equation with an acceptable accuracy, the resolution of the discretization for the domain of interest should be adjusted to the wave number according to the so called rule of thumb [17],





which prescribes a minimum number of elements per wavelength. However, the performance of standard finite difference or finite element methods, such as the classical Galerkin finite element method, deteriorates as *k* increases. This misbehavior, known as pollution of the approximate solution, can only be avoided, with standard methods, after a drastic renement of the discretization. [9] summarized the local truncation error  $|k^2, h_j^2|/k^2$  and the relative phase error  $(k_{\rho}k)|k$  of the most popular methods for solving Helmholtz equation using different size stencils.

Now our aim is to calculate the approximation order of the numerical wavenumber  $\tilde{k}_f$  as k changes, in order to estimate the order of the relative phase error, but first we will examine the order of the error  $|\tilde{k}_f^2 \tilde{\mathcal{K}}_f^2|$  produced by neglecting the dependence of  $\theta$  in taking the average value of  $\tilde{k}_f$  on  $[0.2\pi]$ . We recall that

$$\widetilde{k}_{f}^{2}(k,\theta,h,\varepsilon) = 4E(h,\varepsilon) - 2F(h,\varepsilon)T(k,\theta,h)$$
 (52)

$$k_f^2(k,h,\varepsilon) = 4E(h,\varepsilon) - 6F(h,\varepsilon)J_0(hk)$$
(53)

From the Taylor polynomials of (44) and Bessel function respectively

$$T(k,\theta,h) = 3 - \frac{3}{4}(hk)^2 + \frac{3}{64}(hk)^4 - \frac{10 + \cos(6\theta)}{7680}(hk)^6 + \mathcal{O}(h^8k^8)$$
(54)

and

$$3J_0(hk) = 3 - \frac{3}{4}(hk)^2 + \frac{3}{64}(hk)^4 - \frac{1}{768}(hk)^6 + \mathcal{O}(h^8k^8),$$
(55)

thus  

$$\widetilde{k}_f^2 - k_f^2 = 2F(h,\varepsilon) \left( \frac{\cos(6\theta)}{7680} (hk)^6 + \mathcal{O}(h^8k^8) \right)$$

$$= F(h,\varepsilon) \mathcal{O}(h^6k^6).$$
(56)

From the approximations of  $F(h,\varepsilon)$  given in (34), it is easy to see that

$$F(h,\varepsilon)\mathcal{O}(h^6k^6) = \mathcal{O}(h^4k^6).$$

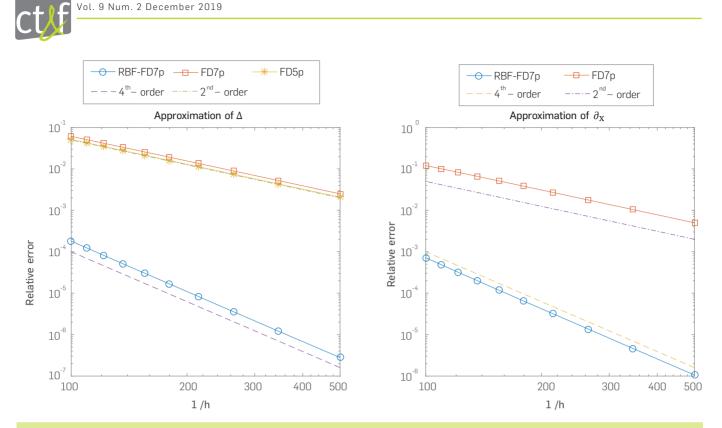
Hence  $\tilde{k}_f^2 - \tilde{k}_f^2 = \mathcal{O}(h^4 k^6)$ . On the other hand, from (34), (55)

$$\widetilde{k}_{f}^{2}(k,\theta,h,\varepsilon) = 4E(h,\varepsilon) - 6F(h,\varepsilon) 
\left(1 - \frac{1}{4}(hk)^{2} + \frac{1}{64}(hk)^{4} + \mathcal{O}(h^{6}k^{6})\right) 
= 4E(h,\varepsilon) - 6F(h,\varepsilon) + \left(\frac{3}{2}h^{2}F(h,\varepsilon)\right)k^{2} 
- \left(\frac{3}{2}h^{2}F(h,\varepsilon)\right)\frac{h^{2}k^{4}}{16} + F(h,\varepsilon)\mathcal{O}(h^{6}k^{6}) 
= k^{2} - \frac{1}{16}h^{2}k^{4} + \varepsilon^{2}h^{2}k^{2} - 2\varepsilon^{4}h^{2} + \mathcal{O}(h^{4}k^{6}) 
+ \mathcal{O}(\varepsilon^{6}h^{4}) + \mathcal{O}(\varepsilon^{2}h^{4}k^{4}) + \mathcal{O}(\varepsilon^{4}h^{4}k^{2}) 
= k^{2}(1 - W).$$
(57)

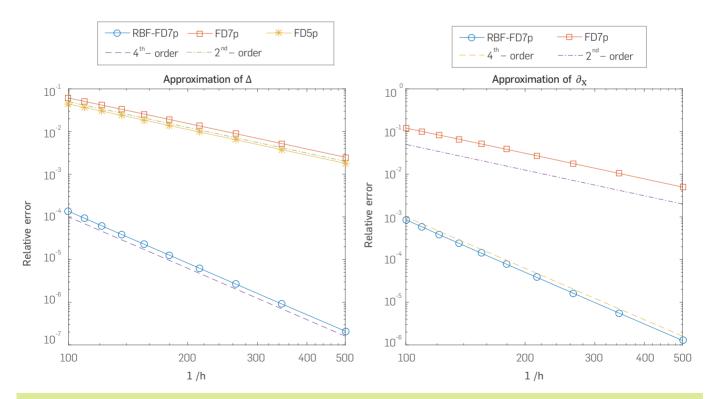
where

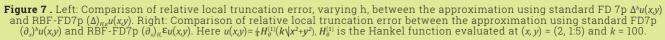
$$W = \frac{1}{16}h^{2}k^{2} - \varepsilon^{2}h^{2} + 2\varepsilon^{4}h^{2}k^{-2} + \mathcal{O}(h^{4}k^{4}) + \mathcal{O}(\varepsilon^{6}h^{4}k^{-2}) + \mathcal{O}(\varepsilon^{2}h^{4}k^{2}) + \mathcal{O}(\varepsilon^{4}h^{4}).$$
(58)

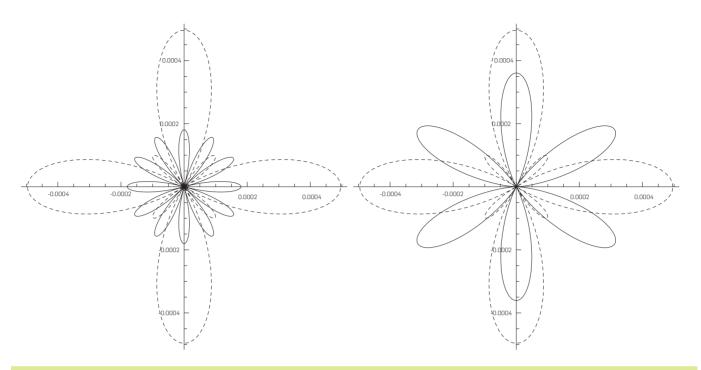
Provided that |W| < 1, the binomial series can be applied, that is  $\sqrt{1 - W} = 1 - \frac{1}{2}W + O(W^2)$ . We take the linear approximation, then



**Figure 6**. Left: Comparison of relative local truncation error, varying *h*, between the approximation using standard FD 7p  $\Delta_{H,0}u(x,y)$  and RBF-FD7p  $\Delta_{H,0}u(x,y)$ . Right: Comparison of relative local truncation error between the approximation using standard FD7p  $(\partial_x)_{H,0}u(x,y)$  and RBF-FD7p  $(\partial_x)_{H,0}u(x,y)$ . Here  $u(x,y)=e^{ik(z\cos\theta+y\sin\theta)}$ , is evaluated at (x,y)=(2.15), with  $\theta=\pi|6$  and k=100.







**Figure 8**. Comparison between the function  $r(\theta) = |\delta k(k_0, \theta, h_o, \varepsilon_{op})|/k_0^2$  (solid line) where  $k_o = 100$  and  $h_o = 0.01$ , and the function  $r(\theta) = J_y(b_o, d_o, e_o; G_o, \theta)/k_0^2$  (dashed line) where  $J_y$  and parameters b0, d0, e0 and G0 were taken from [3]. Left:  $\varepsilon_{op}$  satises  $\delta k(100, 0.001, \varepsilon_{op}) = 0$ . Right:  $\varepsilon_{op} = \arg \min |\delta k(100, \pi/3, 0.01, \varepsilon)|$ 

$$\widetilde{k}_{f}(k,\theta,h,\varepsilon)$$

$$=k-\frac{1}{32}h^{2}k^{3}+\frac{1}{2}\varepsilon^{2}h^{2}k-\varepsilon^{4}h^{2}k^{-1}+\mathcal{O}(h^{4}k^{5})$$

$$+\mathcal{O}(\varepsilon^{6}h^{4}k^{-1})+\mathcal{O}(\varepsilon^{2}h^{4}k^{3})+\mathcal{O}(\varepsilon^{4}h^{4}k).$$
(59)

Hence the relative phase error is given by

$$\frac{k_f - k}{k} = -\frac{1}{32}h^2k^2 + \frac{1}{2}\varepsilon^2h^2 - \varepsilon^4h^2k^{-2} + \mathcal{O}(h^4k^4) + \mathcal{O}(\varepsilon^6h^4k^{-2}) + \mathcal{O}(\varepsilon^2h^4k^2) + \mathcal{O}(\varepsilon^4h^4)$$
(60)

Note that for the limit case  $\varepsilon \rightarrow 0+$ 

$$\widetilde{k}_f(k,\theta,h,0^+) = k + \mathcal{O}(h^2 k^3)$$
(61)

which asserts that, in the limit case, for keeping the order of the error when k is increasing, the mesh size h it must scale at least as  $k-\frac{2}{2}$ , that is  $h = O(k-\frac{3}{2})$ .

# 3. EXPERIMENTAL DEVELOPMENT

Perfectly matched layers (PML) is a technique for simulating solutions of wave phenomena in free-space. The idea is to build an absorbing layer for surrounding the computational domain of interest. The new system possesses the property of generating no reflection, at least in continuous case, at the interface between the free medium and the articial absorbing medium. PML has been introduced by Berenger [1] for Maxwell's equations, and has been widely used for the simulation of time dependent seismic waves as well as Helmholtz-like equations. The analysis necessary to show that the Helmholtz equation with PML (PML-Helmholz) for constant wavenumber is a well-posed problem is treated in [20]. Now a description of PML-Helmholtz equation is given (see [14], [38] for details).

For setting up the PML-Helmholtz equation it is necessary to define some basic functions in a rectangular domain. Let  $\sigma_l \mathbb{R} \rightarrow \mathbb{R}$  be defined by

$$\sigma_{l}(t) = \begin{cases} 0 & \text{for } |t| \leq l \\ \frac{\gamma}{\delta} p\left(\frac{|t|-l}{\delta}\right) & \text{for } l < |t| < l + \delta \\ \frac{\gamma}{\delta} & \text{for } |t| \geq l + \delta \end{cases}$$
(62)

where  $p(0, 1) \rightarrow (0, 1)$  is given by  $p(t) = 1 - e^{\frac{t}{t^2-1}}, \gamma$  and  $\delta$  are constant. By definition  $\sigma \in \mathbb{C}^{\infty}(\mathbb{R})$ . for a fixed positive angular frequency  $\omega$  satisfying  $0 < \frac{\gamma}{\delta} < \omega$ , let dl be the complex-valued function

$$d_l(t) = 1 + i \frac{\sigma_l(t)}{\omega} \tag{63}$$

Consider the open rectangles  $\Omega = (-a,a)\mathbf{x}(-b,b)$  and  $\Omega_{\delta} = (-a-\delta,a+\delta)\mathbf{x}(-b-\delta,b+\delta)$ .

The modified PML-Laplace operator is defined by

$$-\widetilde{\Delta} = -\frac{1}{d_a(x)}\partial_x \left(\frac{1}{d_a(x)}\partial_x\right) - \frac{1}{d_b(y)}\partial_y \left(\frac{1}{d_b(y)}\partial_y\right)$$
(64)



or alternatively

$$-\widetilde{\Delta} = -s_a^2(x)\partial_x^2 - s_b^2(y)\partial_y^2 - s_a(x)s_a'(x)\partial_x - s_b(y)s_b'(y)\partial_y$$
(65)

where  $s_a = \frac{1}{d_a}$  and  $s_b = \frac{1}{d_b}$ . Note that  $-\widetilde{\Delta}$  coincides with  $-\Delta$  in  $\Omega$ . Usually, for making analytic studies this is considered as an unbounded operator on  $L^2(\mathbb{R}^2)$  with domain the Sobolev space  $H^2(\mathbb{R}^2)$  or in a weaker form, as an unbounded operator on  $H^{-1}(\mathbb{R}^2)$  with domain  $H^1(\mathbb{R}^2)$  [38]. However, for numerical implementations in the finite difference approach it is considered a truncated version  $-\widetilde{\Delta}: C^2(\Omega_\delta) \rightarrow C^0(\Omega_\delta)$ . The PML-Laplace operator allows to introduce the PML-Helmholtz equation

$$\begin{bmatrix} -\widetilde{\Delta}u - \omega^2 c^{-2}u &= f, & \text{in } \Omega_{\delta} \\ u &= 0, & \text{on } \partial\Omega_{\delta} \end{bmatrix}$$
 (66)

we will suppose that the function c = c(x) which describe the acoustic or P-wave propagation speed is positive and belongs to  $C^2(\Omega_{\delta})$ .

In order to approximate solutions of equation (66) within the approach of RBF-FD on hexagonal grid, we use the explicit expressions for Gaussian weights, thus by way of the formulas in (33) we have that for x = (x, y) in  $\Omega_{\delta}$  with 7-stencil  $H \subset \Omega_{\delta}$ 

$$\widetilde{\Delta}_{H,\varepsilon}u(\mathbf{x}) = s_a^2(x)(\partial_x^2)_{H,\varepsilon}u(\mathbf{x}) + s_b^2(y)(\partial_y^2)_{H,\varepsilon}u(\mathbf{x}) + s_a(x)s_a'(x)(\partial_x)_{H,\varepsilon}u(\mathbf{x}) + s_b(y)s_b'(y)(\partial_y)_{H,\varepsilon}u(\mathbf{x})$$
(67)

hence by carrying out substitutions

$$\tilde{\Delta}_{H,\varepsilon}u(\mathbf{x}) = \Gamma_{\mathrm{C}}u_{\mathrm{C}} + \Gamma_{\mathrm{e}}u_{\mathrm{e}} + \Gamma_{\mathrm{W}}u_{\mathrm{W}} + \Gamma_{\mathrm{ne}}u_{\mathrm{ne}} + \Gamma_{\mathrm{nw}}u_{\mathrm{nw}} + \Gamma_{\mathrm{se}}u_{\mathrm{se}} + \Gamma_{\mathrm{sw}}u_{\mathrm{sw}}$$
(68)

where weights  $\Gamma$ 's are given by

$$\begin{split} \Gamma_{\rm c}(h,\varepsilon,\mathbf{x}) &= -2E(h,\varepsilon)(s_a^2(x) + s_b^2(y)) \\ \Gamma_{\rm e}(h,\varepsilon,\mathbf{x}) &= F_1(h,\varepsilon)s_a^2(x) + F_2(h,\varepsilon)s_b^2(y) \\ &+ 2D(h,\varepsilon)s_a(x)s_a'(x) \\ \Gamma_{\rm W}(h,\varepsilon,\mathbf{x}) &= F_1(h,\varepsilon)s_a^2(x) + F_2(h,\varepsilon)s_b(y) \\ &- 2D(h,\varepsilon)s_a(x)s_a'(x) \\ \Gamma_{\rm ne}(h,\varepsilon,\mathbf{x}) &= G_1(h,\varepsilon)s_a^2(x) + G_2(h,\varepsilon)s_b^2(y) \\ &+ D(h,\varepsilon)\left(s_a(x)s_a'(x) + \sqrt{3}s_b(y)s_b'(y)\right) \\ \Gamma_{\rm nw}(h,\varepsilon,\mathbf{x}) &= G_1(h,\varepsilon)s_a^2(x) + G_2(h,\varepsilon)s_b^2(y) \qquad (69) \\ &+ D(h,\varepsilon)\left(-s_a(x)s_a'(x) + \sqrt{3}s_b(y)s_b'(y)\right) \\ \Gamma_{\rm SW}(h,\varepsilon,\mathbf{x}) &= G_1(h,\varepsilon)s_a^2(x) + G_2(h,\varepsilon)s_b^2(y) \\ &+ D(h,\varepsilon)\left(-s_a(x)s_a'(x) - \sqrt{3}s_b(y)s_b'(y)\right) \\ \Gamma_{\rm Se}(h,\varepsilon,\mathbf{x}) &= G_1(h,\varepsilon)s_a^2(x) + G_2(h,\varepsilon)s_b^2(y) \end{split}$$

$$+ D(h,\varepsilon) \left( s_a(x) s_a'(x) - \sqrt{3} s_b(y) s_b'(y) \right)$$

and of course,

$$\widetilde{\Delta}_{H,\varepsilon} u(\mathbf{x}) + \omega^2 c^{-2}(\mathbf{x}) u(\mathbf{x}) = (\Gamma_{\rm C} + \omega^2 c^{-2}(\mathbf{x})) u_{\rm C} + \Gamma_{\rm e} u_{\rm e} + \Gamma_{\rm W} u_{\rm W} + \Gamma_{\rm ne} u_{\rm ne} + \Gamma_{\rm nw} u_{\rm nw} + \Gamma_{\rm Se} u_{\rm Se} + \Gamma_{\rm Sw} u_{\rm Sw}$$
(70)

We use the optimal shape parameter  $\varepsilon$  given by (51) to approximate each operator involved in (68). To analyze the approximation  $\widetilde{\Delta}_{H_{\varepsilon}}u(x)$  applied to problem (66) we regard the concept of consistency [32].

**Definition 2.** Given a partial differential equation,  $\mathcal{L}u = f$ , and a finite difference scheme,  $\mathcal{L}_n u = f$  based on a stencil with size h, we say that the finite difference scheme is consistent with the partial differential equation if for any smooth function  $\varphi(\mathbf{x})$ 

$$\mathcal{L}\varphi - \mathcal{L}_h \varphi \to 0 \tag{71}$$

as  $h \rightarrow 0$ , the convergence being pointwise convergence at each point x.

**Proposition 1.** The RBF-FD approximation given in (67) applied to the problem (66) is consistent with the Helmholtz-PML equation and has second order of accuracy.

Proof. We need to show that for an hexagonal 7–stencil H of size h centered at  $x\in\Omega_{\delta}$  and for a given shape parameter  $\varepsilon$ 

$$(-\widetilde{\Delta} - \omega^2 c(\mathbf{x})^{-2}) \varphi(\mathbf{x}) - (-\widetilde{\Delta}_{H,\varepsilon} - \omega^2 c(\mathbf{x})^{-2}) \varphi(\mathbf{x}) = \widetilde{\Delta}_{H,\varepsilon} \varphi(\mathbf{x}) - \widetilde{\Delta} \varphi(\mathbf{x})$$
(72)

converges pointwise to  $\theta$  as  $h \rightarrow \theta$ , where  $\varphi$  is a smooth function. The proof will be completed by simply inspecting the estimates in (38) used in (67), once performed these substitutions we can observe that

$$|\widetilde{\Delta}_{H,\varepsilon}\varphi(\mathbf{x}) - \widetilde{\Delta}\varphi(\mathbf{x})| = \mathcal{O}(\varepsilon^2 h^2),$$
(73)

which means that the accuracy of the approximation is  $\mathcal{O}(\varepsilon^2 h^2)$ , as  $\varepsilon h$  tends to 0.

#### **DISCRETE FORMULATION**

We will perform the discretization of the problem (66) on an hexagonal grid G which represents to  $\Omega_{\delta}$ . The optimal shape parameter at every stencil will be taken in accord to its dependence of the local wavenumber  $k(x) = \omega^2 c^{-2}(x)$ , so  $\varepsilon_{op}(x) = \varepsilon_{op}(k(x),h)$ , where is used (51). For the sake of a more simplied notation, we denote similarly for the others weights formulas in (69).

$$\Gamma_{\rm C}(\mathbf{x}) = \Gamma_{\rm C}(h, \varepsilon_{op}(\mathbf{x}), \mathbf{x}), \tag{74}$$

We begin by considering two rectangular grids given by

$$G^{1} = \left\{ \left(mh, n\sqrt{3}h\right) : m = 0, \dots, \\ N_{x} - 1; n = 0, \dots, \frac{N_{z} - 1}{2} \right\}$$

$$G^{2} = \left\{ \left(\left(m - \frac{1}{2}\right)h, \left(n + \frac{1}{2}\right)\sqrt{3}h\right) \\ : m = 0, \dots, N_{x}; n = 0, \dots, \frac{N_{z} - 3}{2} \right\}$$
(75)

where  $N_z$  is a positive odd integer. Note that the grid  $G = G^1 \cup G^2$  is formed by hexagonal stencils. We denote  $x_{nm}^I = (mh, n\sqrt{3}h)$  and  $x_{nm}^2 = ((m - \frac{1}{2})h, (n + \frac{1}{2})\sqrt{3}h)$  for points in  $G^1$  and  $G^2$  respectively; also we denote  $u_{nn}^I = u(x_{nn}^I)$  for  $x_{nn}^I \in Gj$ , j = 1, 2. The homogeneous Dirichlet boundary condition it is expressed in terms of the boundary points of the grid G, by doing

$$u_{m,0}^{1} = u_{m,\frac{Nz-1}{2}}^{1} = u_{N_{x}-1,n}^{1} = u_{0,n}^{1}$$
  
= 0; for  $m = 0, \dots, N_{x} - 1; n = 0, \dots, \frac{N_{z} - 1}{2}$  (76)

$$u_{0,n}^{2} = u_{N_{x,n}}^{2} = 0; \text{ for } n = 0, \dots, \frac{N_{z} - 3}{2}$$
(77) Thus, for points in the grid G1, the equation (68) has the explicit form  $\Gamma_{SW} u_{m,n-1}^{2} + \Gamma_{SE} u_{m+1,n-1}^{2} + \Gamma_{W} u_{m-1,n}^{1} + \Gamma_{C} u_{m,n}^{1} + \Gamma_{E} u_{m+1,n}^{1} + \Gamma_{RW} u_{m-1,n}^{2} + \Gamma_{E} u_{m+1,n}^{2} + \Gamma_{RW} u_{m+1,n}^{2} + \Gamma_{RW} u_{m+1,n}^{2} + \Gamma_{RW} u_{m+1,n}^{2} + \Gamma_{RW} u_{m-1,n}^{2} + \Gamma_{RW} u_{m-$ 

is a block matrix whose blocks are given by

$$\mathbf{C}_{n}^{1} = \begin{pmatrix} \Gamma_{c}(\mathbf{x}_{1,n}^{1}) & \Gamma_{c}(\mathbf{x}_{2,n}^{1}) & \Gamma_{e}(\mathbf{x}_{3,n}^{1}) & 0 & 0 \\ \Gamma_{W}(\mathbf{x}_{1,n}^{1}) & \Gamma_{c}(\mathbf{x}_{2,n}^{1}) & \Gamma_{e}(\mathbf{x}_{3,n}^{1}) & 0 & 0 \\ 0 & \Gamma_{W}(\mathbf{x}_{2,n}^{1}) & \Gamma_{c}(\mathbf{x}_{3,n}^{1}) & 0 & 0 \\ \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \Gamma_{c}(\mathbf{x}_{N-3,n}^{1}) & \Gamma_{e}(\mathbf{x}_{N-2,n}^{1}) \\ 0 & 0 & 0 & \Gamma_{W}(\mathbf{x}_{N-3}^{1}) & \Gamma_{c}(\mathbf{x}_{N-2,n}^{1}) \\ 0 & 0 & 0 & \Gamma_{W}(\mathbf{x}_{N-3}^{2}) & \Gamma_{c}(\mathbf{x}_{2,n}^{1}) \\ \Gamma_{W}(\mathbf{x}_{1,n}^{2}) & \Gamma_{c}(\mathbf{x}_{2,n}^{2}) & \Gamma_{c}(\mathbf{x}_{3,n}^{2}) & 0 & 0 \\ 0 & \Gamma_{W}(\mathbf{x}_{2,n}^{2}) & \Gamma_{c}(\mathbf{x}_{3,n}^{2}) & 0 & 0 \\ 0 & \Gamma_{W}(\mathbf{x}_{2,n}^{2}) & \Gamma_{c}(\mathbf{x}_{3,n}^{2}) & 0 & 0 \\ \vdots & \ddots & \\ 0 & 0 & 0 & \Gamma_{C}(\mathbf{x}_{N-2,n}^{2}) & \Gamma_{c}(\mathbf{x}_{N-1,n}^{2}) \\ 0 & 0 & 0 & \Gamma_{W}(\mathbf{x}_{N-2,n}^{2}) & \Gamma_{c}(\mathbf{x}_{N-1,n}^{2}) \end{pmatrix}$$

$$\mathbf{N}_{n}^{1} = \begin{pmatrix} \Gamma_{ne}(\mathbf{x}_{1,n}^{1}) & 0 & \cdots & 0 & 0 \\ \Gamma_{nw}(\mathbf{x}_{1,n}^{1}) & \Gamma_{ne}(\mathbf{x}_{1,n}^{1}) & 0 & \cdots & 0 & 0 \\ \Gamma_{nw}(\mathbf{x}_{1,n}^{1}) & \Gamma_{ne}(\mathbf{x}_{1,n}^{2}) & 0 & 0 \\ \vdots & \ddots & \\ 0 & 0 & \Gamma_{nw}(\mathbf{x}_{1,n}^{1}) & 0 & \cdots & 0 & 0 \\ \vdots & \ddots & \vdots \\ 0 & 0 & \Gamma_{nw}(\mathbf{x}_{1,n}^{1}) & 0 & \cdots & 0 & 0 \\ \vdots & \ddots & \vdots \\ 0 & 0 & \Gamma_{nw}(\mathbf{x}_{1,n}^{1}) & 0 & \cdots & 0 & 0 \\ \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \Gamma_{nw}(\mathbf{x}_{1,n-2,n}^{1}) & \Gamma_{ne}(\mathbf{x}_{1,n-2,n}^{1}) \\ 0 & 0 & \cdots & 0 & \Gamma_{nw}(\mathbf{x}_{1,n-2,n}^{1}) \end{pmatrix}$$

$$(83)$$

$$\mathbf{N}_{n}^{2} = \begin{pmatrix} \Gamma_{\mathrm{nw}}(\mathbf{x}_{1,n}^{2}) \ \Gamma_{\mathrm{ne}}(\mathbf{x}_{2,n}^{2}) & 0 & \cdots & 0 & 0 \\ 0 & \Gamma_{\mathrm{nw}}(\mathbf{x}_{2,n}^{2}) \Gamma_{\mathrm{ne}}(\mathbf{x}_{3,n}^{2}) & 0 & 0 \\ \vdots & & \ddots & & \vdots \\ 0 & 0 & 0 & \Gamma_{\mathrm{ne}}(\mathbf{x}_{N_{x}-2,n}^{2}) & 0 \\ 0 & 0 & 0 & \Gamma_{\mathrm{ne}}(\mathbf{x}_{N_{x}-2,n}^{2}) \ \Gamma_{\mathrm{ne}}(\mathbf{x}_{N_{x}-1,n}^{2}) \end{pmatrix}$$
(84)

$$\mathbf{S}_{n}^{1} = \begin{pmatrix} \Gamma_{\rm Se}(\mathbf{x}_{1,n}^{1}) & 0 & \cdots & 0 & 0 \\ \Gamma_{\rm SW}(\mathbf{x}_{1,n}^{1}) & \Gamma_{\rm Se}(\mathbf{x}_{2,n}^{1}) & 0 & 0 \\ 0 & \Gamma_{\rm SW}(\mathbf{x}_{2,n}^{1}) & 0 & 0 \\ \vdots & \ddots & \vdots & \ddots & \vdots \end{pmatrix}$$
(85)

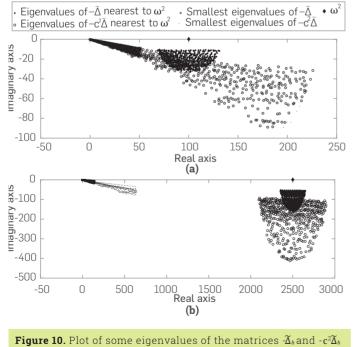
$$\mathbf{S}_{n}^{2} = \begin{pmatrix} 0 & 0 & \Gamma_{SW}(\mathbf{x}_{N_{x}-3,n}^{*}) & \Gamma_{Se}(\mathbf{x}_{N_{x}-2,n}^{*}) \\ 0 & 0 & \cdots & 0 & \Gamma_{SW}(\mathbf{x}_{N_{x}-2,n}^{1}) \\ \Gamma_{SW}(\mathbf{x}_{1,n}^{2}) & \Gamma_{Se}(\mathbf{x}_{2,n}^{2}) & 0 & \cdots & 0 & 0 \\ 0 & \Gamma_{SW}(\mathbf{x}_{2,n}^{2}) & \Gamma_{Se}(\mathbf{x}_{3,n}^{2}) & 0 & 0 \\ \vdots & & \ddots & \vdots \\ 0 & 0 & 0 & \Gamma_{Se}(\mathbf{x}_{N_{x}-2,n}^{2}) & 0 \\ 0 & 0 & 0 & \Gamma_{Se}(\mathbf{x}_{N_{x}-2,n}^{2}) & \Gamma_{Se}(\mathbf{x}_{N_{x}-1,n}^{2}) \end{pmatrix}$$
(86)

$$\mathbf{U}_{h} = \left(u_{1,1}^{2} \cdots u_{N_{x}-1,1}^{2} \ u_{1,1}^{1} \cdots u_{N_{x}-2,1}^{1} \ u_{1,2}^{2} \cdots u_{N_{x}-1,2}^{2} \cdots u_{1,\frac{N_{z}-5}{2}}^{1} \cdots \right)$$
(87)

$$u_{N_x-2,\frac{N_z-5}{2}}^1 u_{1,\frac{N_z-3}{2}}^2 \cdots u_{N_x-1,\frac{N_z-3}{2}}^2 \right)$$
(88)

$$\mathbf{f} = \left( f_{1,1}^2 \cdots f_{N_x-1,1}^2 f_{1,1}^1 \cdots f_{N_x-2,1}^1 f_{1,2}^2 \cdots f_{N_x-1,2}^2 \cdots f_{1,\frac{N_z-5}{2}}^1 \cdots \right)$$
(89)

$$f_{N_x-2,\frac{N_z-5}{2}}^1 f_{1,\frac{N_z-3}{2}}^2 \cdots f_{N_x-1,\frac{N_z-3}{2}}^2$$
(90)



(Top) for  $\omega = 10$  and h = 0.1497, (below)  $\omega = 50$  and h = 0.0300. c represents the Marmousi model extended to PML.

The sparsity pattern of the matrix  $\tilde{\Delta}_h$  is depicted in **Figure 11**.

Remark 3. If  $L_h = (-\widetilde{\Delta}_h - \omega^2 c^2)$ , the linear system  $L_h U_h = \mathbf{f}$  can be solved if only if  $\omega^2 \notin \sum (-c^2 \widetilde{\Delta}_h)$ . Here  $\sum (\mathbf{A})$  denotes the spectrum of  $\mathbf{A}$ . Is pending for a further work to prove that  $\omega^2 \notin \sum (-c^2 \widetilde{\Delta}_h)$ . Figure 10 shows a plot of some eigenvalues of the matrices  $-\widetilde{\Delta}_h$  and  $-c^2 \widetilde{\Delta}_h$ , where c represents the Marmousi model extended to PML. The eigenvalues plotted are 500 nearest to  $\omega^2$  and 500 nearest to 0.

### 4. RESULTS

#### EXACT IMPEDANCE BOUNDARY CONDITIONS

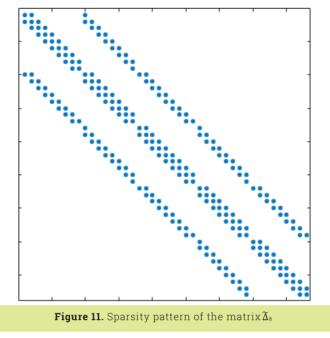
$$\begin{array}{rcl} -\Delta u - k^2 u &=& f, \text{ in } \Omega\\ ik\mathcal{B}u + \frac{\partial u}{\partial n} &=& g, \text{ on } \partial\Omega \end{array} \tag{91}$$

#### PLANE WAVES

With  $\Omega = (0,1) \times (0,1)$ , k = 500,  $k_x = k \cos \theta$ ,  $k_z = k \sin \theta$ . By using  $2\pi$  points per wavelength, *i.e.* kh = 1, we solve (91)

with g given by 
$$g(x, z) = \begin{cases} i(k - k_z)e^{ik_x x}, & \text{if } z = 0\\ i(k + k_x)e^{i(k_x + k_z z)} & \text{if } x = 1\\ i(k + k_z)e^{i(k_x x + k_z)} & \text{if } z = 1\\ i(k - k_x)e^{ik_z z} & \text{if } x = 0 \end{cases}$$

0



The results can be seen in Figure13.

#### SPHERICAL WAVEFRONT

For this example it is chosen  $N_g = 6$  for solving the problem (91) with  $\Omega = (0, 1) \times (0, 1)$ . Where g is the boundary data satisfying the exact solutions

$$u_1(x,y) = H_0^{(1)}(k\sqrt{(x-2)^2 + (y-2)^2}), \qquad (92)$$

and

$$u_{2}(x,y) = \sqrt{k}H_{0}^{(1)}(k\sqrt{(x+20)^{2}+(y+20)^{2}}) + 2\sqrt{k}H_{0}^{(1)}(k\sqrt{(x-20)^{2}+(y-20)^{2}}) + 0.5\sqrt{k}H_{0}^{(1)}(k\sqrt{(x+20)^{2}+(y-20)^{2}}) - \sqrt{k}H_{0}^{(1)}(k\sqrt{(x-20)^{2}+(y+20)^{2}}),$$
(93)

with constant *k*. These solutions corresponds to a single source and four sources, respectively, outside the domain  $\Omega$ . The relevancy of taking  $u_2$  can be seen in [13]. Discretizations were made with  $N_g = 6$  nodes per wavelength at each value of the wavenumber k. Within an hexagonal grid we take 7–stencils at inner nodes and 22–stencils at boundary nodes.

<b>Table 1.</b> Error of the approximated solutions $\tilde{u}_1$ and $\tilde{u}_2$ for the	
problem (91) with source outside the domain, whose exact	
solution $u_1$ and $u_2$ are given in (92) and (93).	

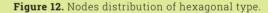
k	50	100	200	400	800
2π~h	30	600	1200	2400	4200
Ν	2691	10604	42214	168847	674614
N <sub>nh</sub>	608	1379	3441	5460	10106
N <sub>b</sub>	206	412	822	1646	3290
$  u_c - \widetilde{u}_c  _{L^2}$	1.28e-03	9.84e-04	8.15e-04	6.11e-04	6.11e-04
$  u_c - \widetilde{u}_c  _{L^2} /   u_c  _{L^2}$	1.78e-02	1.91e-02	2.24e-02	2.38e-02	3.36e-02
$  u_c - \widetilde{u}_c  _{L^2}$	7.27e-03	7.15e-03	7.41e-03	6.07e-03	6.40e-03
$  u_c - \widetilde{u}_c  _{L^2} /   u_c  _{L^2}$	2.09e-02	2.04e-02	2.11e-02	1.73e-02	1.82e-02

#### **INSIDE SOURCE PROBLEM**

0 0

Now we test our method with sources inside the domain. The aim is to get the truncated solution of the problem

$$u_{c}(x,y) = \frac{i}{4}H_{0}^{(1)}(k\sqrt{(x-0.5)^{2} + (y-0.5)^{2}}), \quad (94)$$



$$u_M(x,y) = \frac{i}{4} H_0^{(1)} \left( k \sqrt{(x-0.5)^2 + (y-0.5)^2} \right) + \frac{i}{4} H_0^{(1)} \left( k \sqrt{(x-0.25)^2 + (y-0.25)^2} \right),$$
(95)

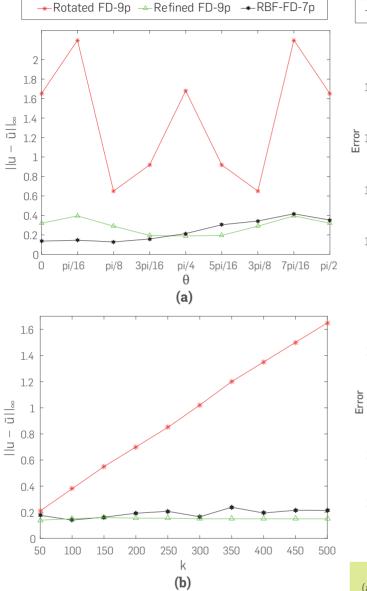
**Table 2.** Error of the approximated solutions  $\tilde{u}_1$  and  $\tilde{u}_2$  for the problem (91) with source *f* inside the domain, g = 0 and  $\mathcal{B}$  the third order Pade approximation. Exact solutions  $u_c$  and  $u_M$  are given in (94) and (95).

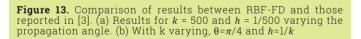
k	50	100	200	400	800
2π~h	30	600	1200	2400	4200
Ν	2691	10604	42214	168847	674614
N <sub>nh</sub>	608	1379	3441	5460	10106
Nb	206	412	822	1646	3290
$  u_c - \widetilde{u}_c  _{L^2}$	3.13e-03	1.69e-03	1.97e-03	1.27e-03	8.70e-04
$  u_c - \widetilde{u}_c  _{L^2} /   u_c  _{L^2}$	6.45e-02	4.91e-02	8.00e-02	7.28e-02	7.03e-02
им-Ũм   <sub>L<sup>2</sup></sub>	4.15e-03	2.72e-03	2.38e-03	1.73e-03	1.27e-03
U <sub>M</sub> -Ũ <sub>M</sub>    <sub>1</sub> 2/  U <sub>M</sub>    <sub>1</sub> 2	6.54e-02	6.10e-02	6.83e-02	7.39e-02	7.34e-02

#### WITH PML

In this test we solve the problem of nd the truncated Green's function for  $-\tilde{\Delta}_h - k^2$  in  $\Omega = (0, 1) \times (0, 1)$ , that is the solution of

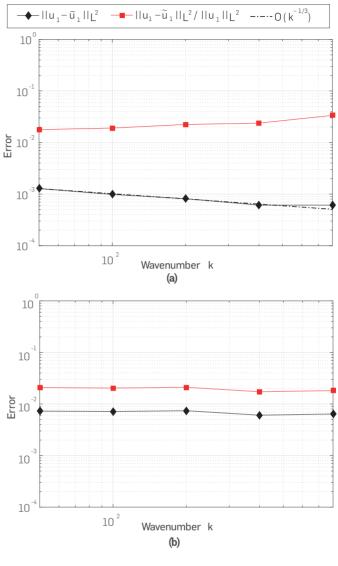
$$\begin{split} &(-\widetilde{\Delta}-k^2)u(x,z) = \widetilde{\delta}(x-0.25,z-5\times 10^{-4}) \\ &+ \widetilde{\delta}(x-0.5,z-5\times 10^{-4}) + \widetilde{\delta}(x-0.75,z-5\times 10^{-4}). \end{split}$$

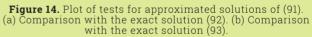




The solution of this problem is

$$u(x,y) = H_0^{(1)} (k\sqrt{(x-0.25)^2 + (z-5\times 10^{-4})^2}) + H_0^{(1)} (k\sqrt{(x-0.5)^2 + (z-5\times 10^{-4})^2}) + H_0^{(1)} (k\sqrt{(x-0.75)^2 + (z-5\times 10^{-4})^2})$$
(96)





where  $H_0^{(1)}$  is the Hankel function. For this test we use  $2\pi$  points per wavelength (*i.e.* kh = 1). Some results with a qualitative analysis of error are shown in Figure 15.

#### 2004 BP MODEL

In this example we calculate the truncated Green function for the well known BP model in frequencies 15Hz and 40Hz. The results are shown in **Figure 16**.

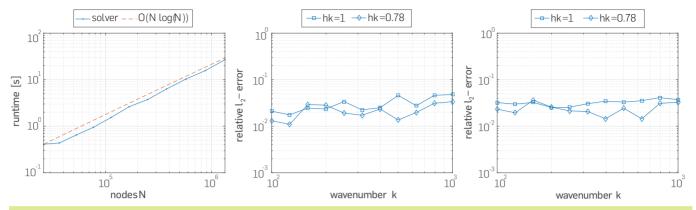


Figure 15. (Left) Runtime for solving the system  $L_h U_h$  = f by LU factorization. (Center) Near boundary sources. (Right) Centered sources.

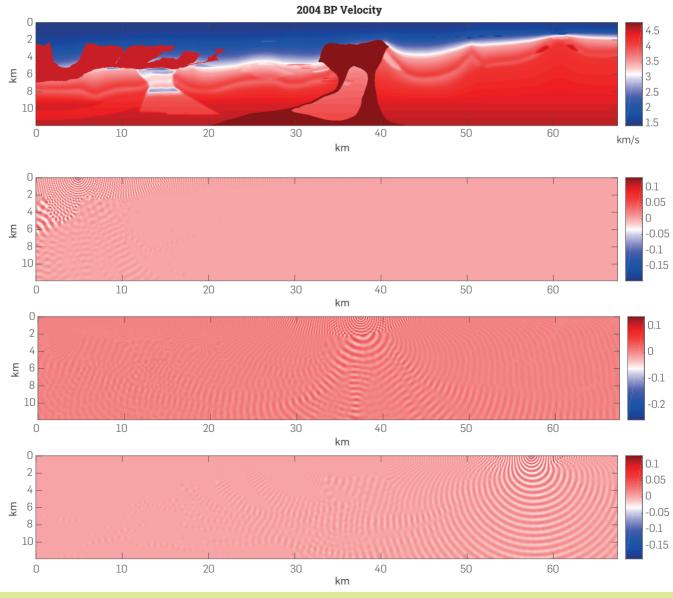


Figure 16. Top: 2004 BP velocity-analysis Benchmark. Bottom plots: Real part of the wave eld at 6Hz with dierent positions of the source.

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## **CONCLUSIONS**

In this paper we have developed and applied a RBF-FD method on hexagonal grids that solves the Helmholtz equation for a wide range of values of the wavenumber *k*. We focus in obtaining accurate local approximations of the partial derivative operators by using a few degrees of freedom applied to solutions of Helmholtz equation.

We obtain closed formulas for the weights in terms of  $\varepsilon$ , *h* showing that the standard finite difference method is the limit of our RBF-FD scheme, this happens when the wide of the Gaussian RBF, controlled by the shape parameter  $\varepsilon$ , tends to be flat. The method obtains banded matrices; such that the respective linear systems can be solved efficiently by direct solvers.

We have developed a simple solver that is compared with 5-point Finite difference method and accurate compared with 9-point FD method. By local interpolation we found good shape parameters to approximate derivatives of solutions of the 2D Helmholtz equation. Within the RBF-FD framework are approximated the solutions of the PMLHelmholtz equation on hexagonal grids with optimal values of the shape parameter.

Our method has been compared against classical and well-known methods, showing equal and sometimes better performance. with this we have validated numerically that our solver suer less of the pollution effects. The advantages and drawbacks of the RBF-FD method applied to Helmholtz equation can be seen at **Table 3**.

## **ACKNOWLEDGEMENTS**

Table 3. Advantages and	drawbacks of RBF-FD method
-------------------------	----------------------------

Strength	Drawbacks
It doesn't require a mesh. Rather than discretizing space, the solutions are approximated by a set of basis functions. The method only needs to use collocation points in the domain and on the poorder	The resulting linear equation system may have a bad conditioning. This problem has been treated with different matrix conditioning methods
Simpler numerical implementa- tion and lower computational cost	Lower quality approach near the border.
You can easily increase the support domain, change basis, adopt nodal densities or local refinement.	It requires tuning parameters. Strongly depends on the proper value of shape parameter, whose optimal value is not known a priori
Easily adaptable to irregular domains or higher dimensions.	It is difficult to find formulas for error bounds in interpolation
t is efficient in terms of computa- tional time and can be readily parallelized	If the number of nodes changes, it is necessary to change the shape parameter accordingly

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